

Genealogy of the
N-particle branching random walk
with polynomial tails

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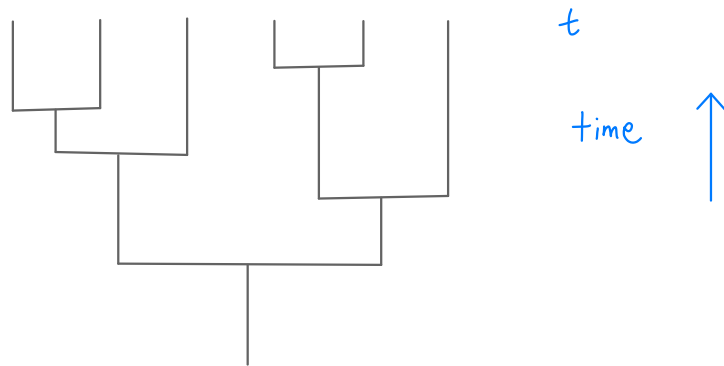
Joint work with Matt Roberts and Zsófia Talyigás

Branching-selection systems

- Particle systems: particles branch (produce offspring) and move in space
killing rule keeps total number of particles constant.
- Toy models for a population under selection.
Location of a particle (= individual) represents its evolutionary fitness.
- Introduced by Brunet and Derrida in 1990s.
Recent results and open conjectures about long-term behaviour.

Genealogy:

Coalescent process



N-particle branching random walk (N-BRW)

Discrete-time branching-selection system.

N particles with locations in \mathbb{R} at each timestep.

Let X be a real-valued random variable (jump distribution).

At each time $n \in \mathbb{N}_0$, each particle has two offspring.

Each of the $2N$ offspring particles makes an independent jump from its parent's location, with the same law as X .

The N rightmost particles (of the $2N$ offspring particles) form the population at time $n+1$.



Notation: $X_1^{(N)}(n) \leq X_2^{(N)}(n) \leq \dots \leq X_N^{(N)}(n)$ ordered particle positions at time n .

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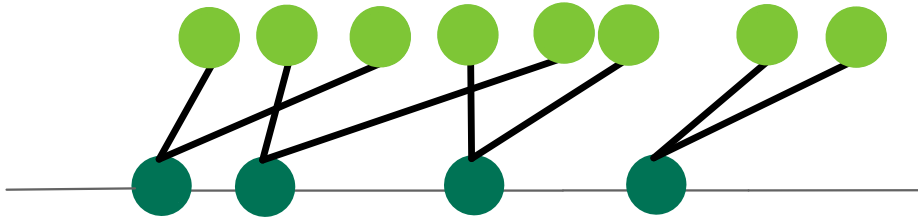
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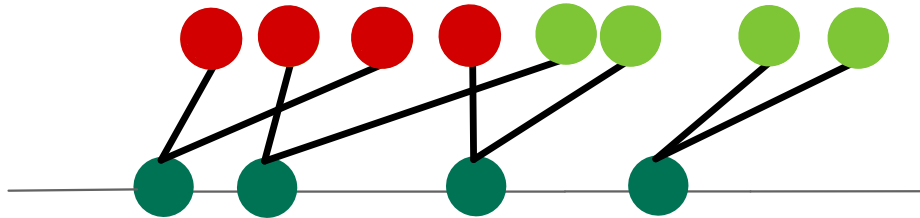
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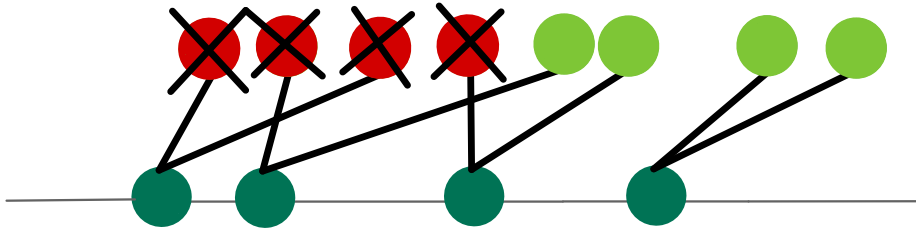
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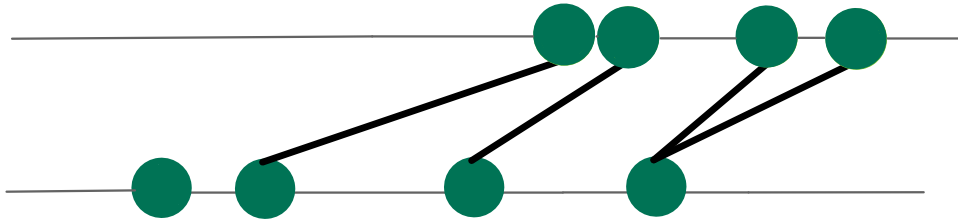
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Light-tailed jump distribution

$$\mathbb{P}(X > x) \leq e^{-cx}, \quad c > 0$$

Asymptotic speed

If $\mathbb{E}[X] < \infty$ then $\exists v_N \in (0, \infty)$ s.t.

$$\lim_{n \rightarrow \infty} \frac{X_N^{(N)}(n)}{n} = v_N = \lim_{n \rightarrow \infty} \frac{X_1^{(N)}(n)}{n} \quad \text{a.s. and in } L^1.$$

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Theorem (Bérard and Guéré 2010) If $\mathbb{E}[e^{\lambda X}] < \infty$ for some $\lambda > 0$ (+technical assumptions) then $\lim_{N \rightarrow \infty} v_N = v_\infty$ exists and $v_\infty - v_N \sim c(\log N)^{-2}$ as $N \rightarrow \infty$.

Conjectured by Brunet + Derrida 1997. Related result for Fisher-KPP equation with noise (Mueller, Mytnik, Quastel 2009)

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Genealogy

Sample k particles from the N particles and trace their ancestry backwards in time \rightarrow coalescent process.

Conjecture (Brunet, Derrida, Mueller, Munier)

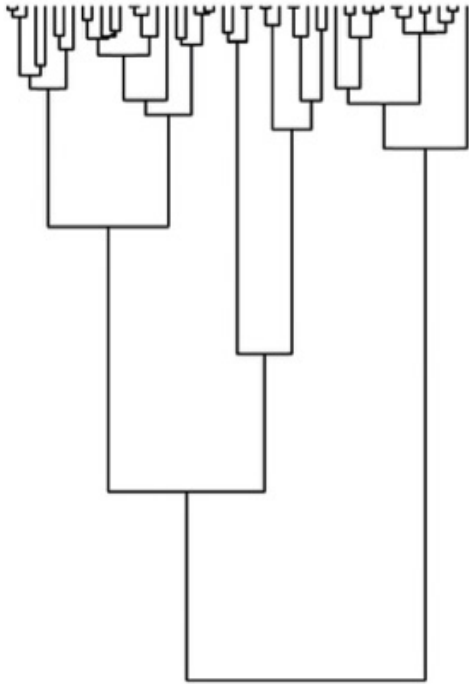
If X is light-tailed then the genealogy of a sample on a $(\log N)^3$ timescale converges to a Bolthausen-Sznitman coalescent as $N \rightarrow \infty$.

See Berestycki, Berestycki, Schweinsberg.

Coalescent processes

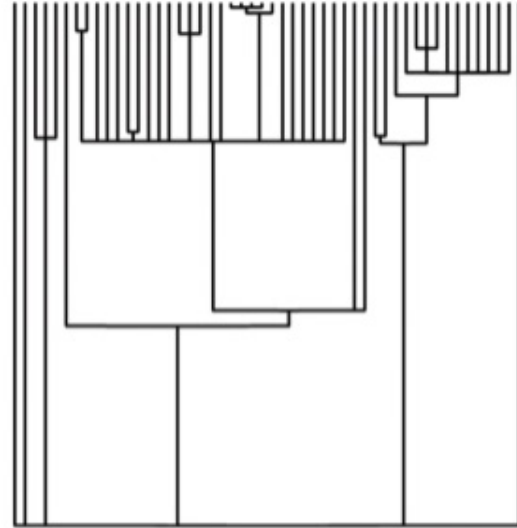
Kingman's coalescent

Neutral population: choose particles to kill uniformly at random in each generation.



Bolthausen-Sznitman coalescent

Population under selection.



Thanks to Götz Kersting
and Anton Wakolbinger

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N-BRW with heavy-tailed jump distribution

Suppose $P(X > x) \sim x^{-\alpha}$ as $x \rightarrow \infty$, for some $\alpha > 0$.

Asymptotic speed

Theorem (Bérard and Maillard 2014)

If $E[X] < \infty$, $\lim_{n \rightarrow \infty} \frac{X_N^{(N)}(n)}{n} = v_N$ where $v_N \sim c_\alpha N^{1/\alpha} (\log N)^{1/\alpha - 1}$ as $N \rightarrow \infty$.

If $E[X] = \infty$, cloud of particles accelerates.

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Genealogy

Conjecture (Bérard and Maillard)

The genealogy on a $\log N$ timescale is approximately given by a star-shaped coalescent when N is large.

Time and space scales

$$\text{Let } \mathbb{P}(X > x) = \frac{1}{h(x)} \text{ for } x \geq 0.$$

Assume h is regularly varying with index $\alpha > 0$

$$\text{i.e. for any } y > 0, \quad \frac{h(xy)}{h(x)} \longrightarrow y^\alpha \text{ as } x \rightarrow \infty.$$

and $\mathbb{P}(X \geq 0) = 1$ (no negative jumps).

$$\text{e.g. } h(x) = x^\alpha \text{ for } x \geq 1$$

$$\text{Let } \ell_N = \lceil \log_2 N \rceil \text{ time scale}$$

$$\text{Let } a_N = h^{-1}(2N\ell_N), \quad \text{where } h^{-1}(x) := \inf \{y \geq 0 : h(y) > x\}. \quad \text{space scale}$$

$$\begin{aligned} \mathbb{E} \left[\begin{array}{l} \# \text{ jumps of size } > c_1 a_N \text{ in} \\ \text{a time interval of length } c_2 \ell_N \end{array} \right] &= 2N \cdot c_2 \ell_N \mathbb{P}(X > c_1 a_N) \\ &= \frac{2N c_2 \ell_N}{h(c_1 a_N)} \sim \frac{2N c_2 \ell_N}{c_1^\alpha 2N \ell_N} = \frac{c_2}{c_1^\alpha} \\ &\quad \text{as } N \rightarrow \infty. \end{aligned}$$

Main result

w.h.p. = with probability $\rightarrow 1$ as $N \rightarrow \infty$.

Theorem (P., Roberts, Talyigás 2021)

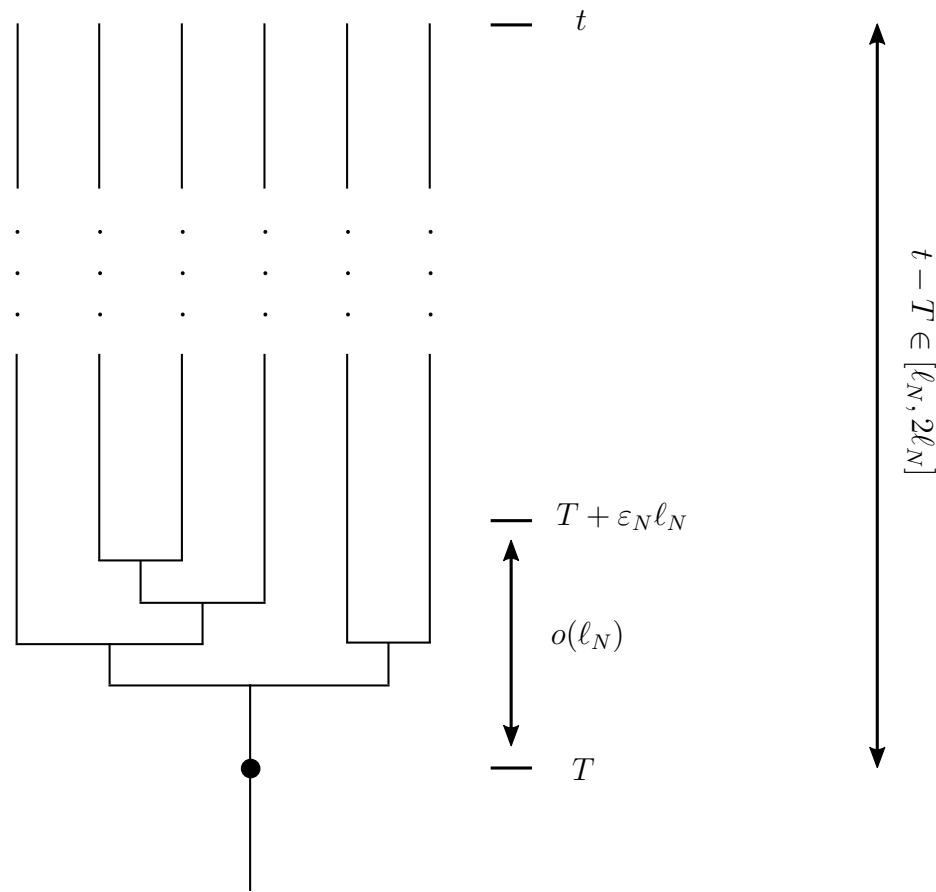
For $\eta > 0$, $k \in \mathbb{N}$ and $t > 4\ell_N$, the following occurs w.h.p.:

- **Spatial distribution:** At time t , there are $N - o(N)$ particles in

$$[X_i^{(N)}(t), X_i^{(N)}(t) + \eta a_N].$$

- **Genealogy:** The genealogy on an ℓ_N -timescale is asymptotically given by a star-shaped coalescent.

i.e. $\exists T \in [t - 2\ell_N, t - \ell_N]$ s.t. w.h.p., for a uniform sample of k particles at time t , every particle is descended from the rightmost particle at time T and no pair of particles in the sample has a common ancestor after time $T + \varepsilon_N \ell_N$, for any $(\varepsilon_N)_N$ with $\varepsilon_N \rightarrow 0$ and $\varepsilon_N \ell_N \rightarrow \infty$ as $N \rightarrow \infty$.



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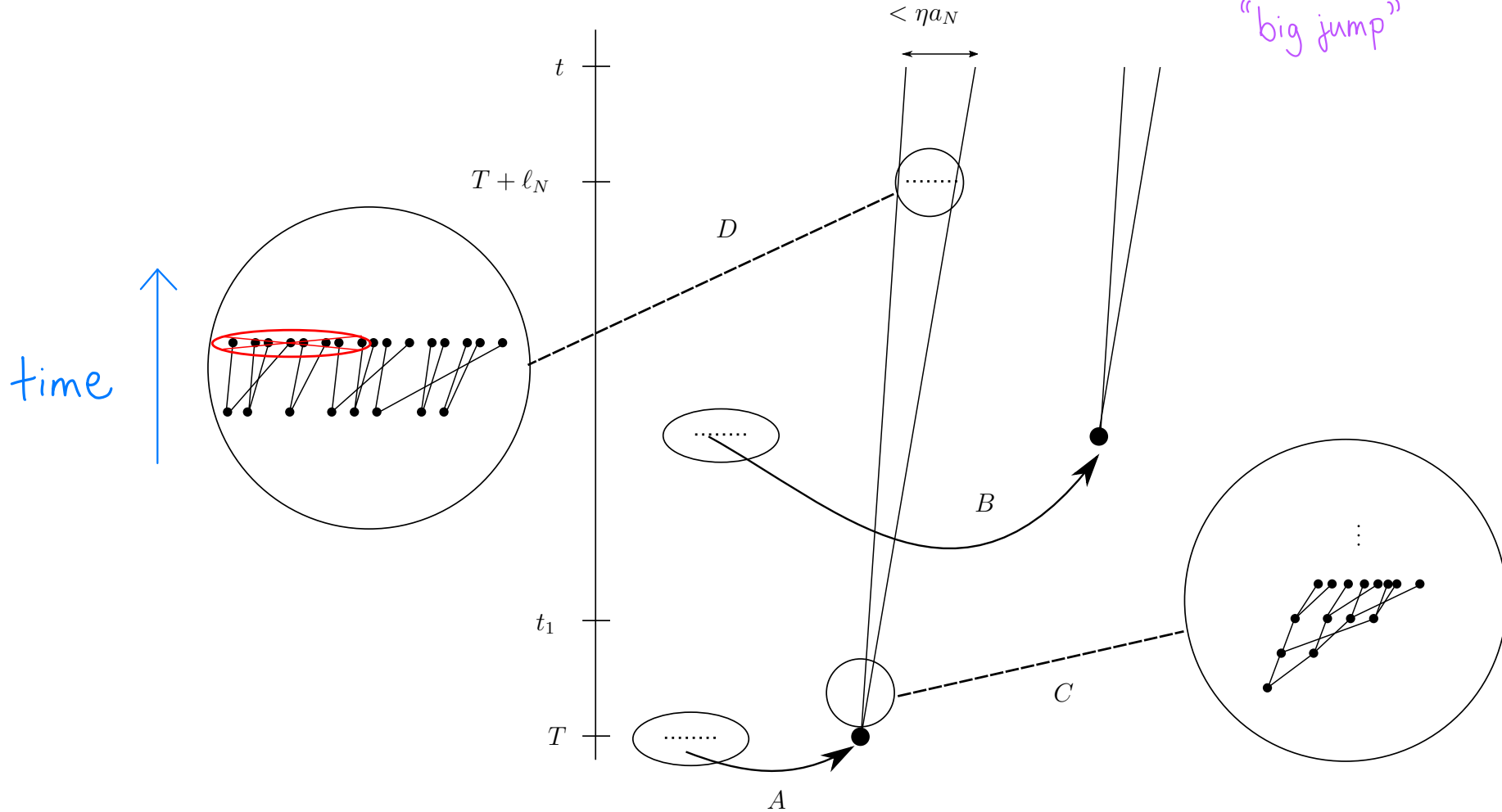
Proof heuristics

$$t_1 := t - \ell_N.$$

$\rho > 0$ small constant.

Let $T =$ last time before time t_1 when a particle makes a jump $\geq \rho a_N$ and takes the lead.

"big jump"

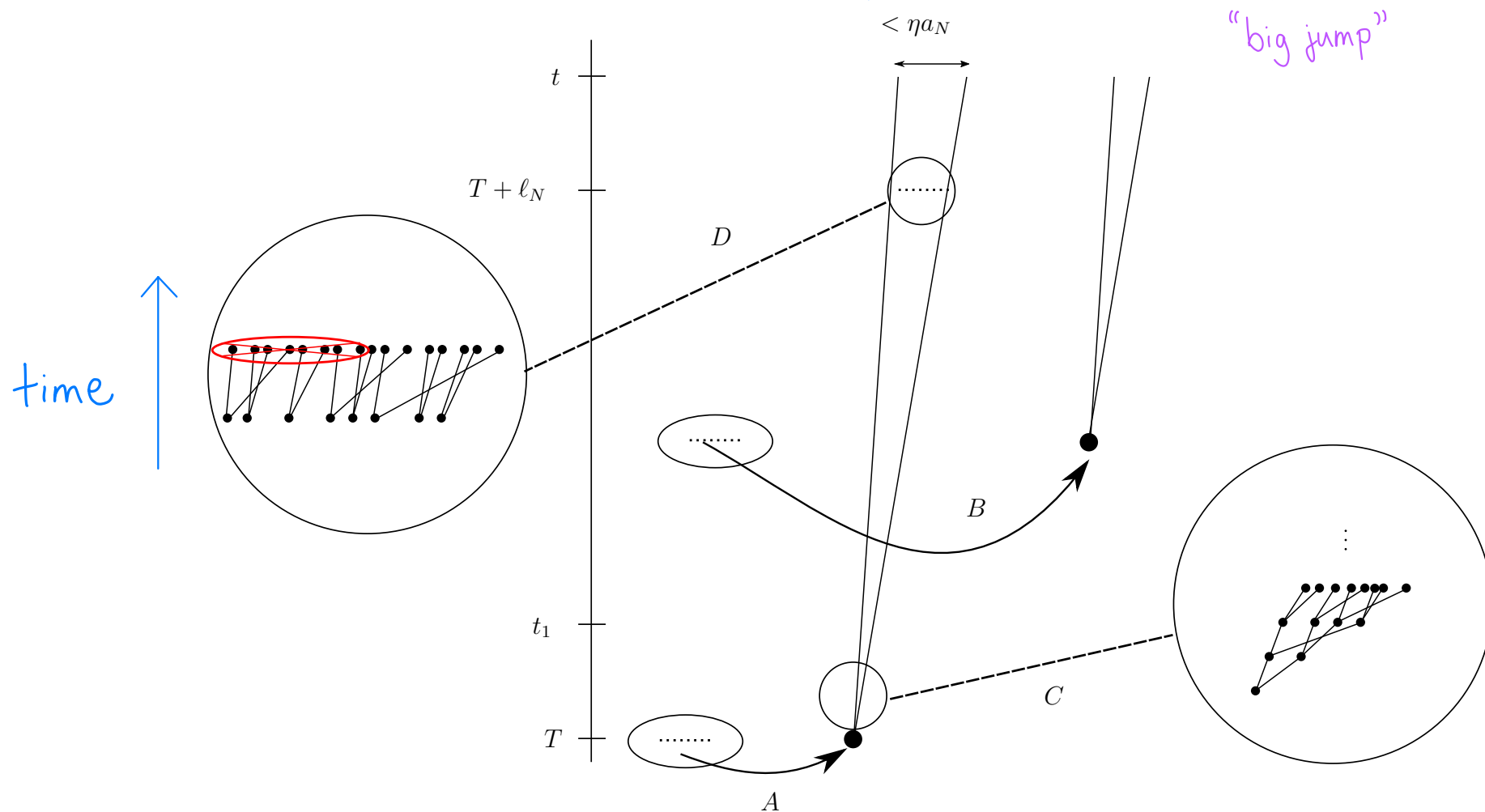


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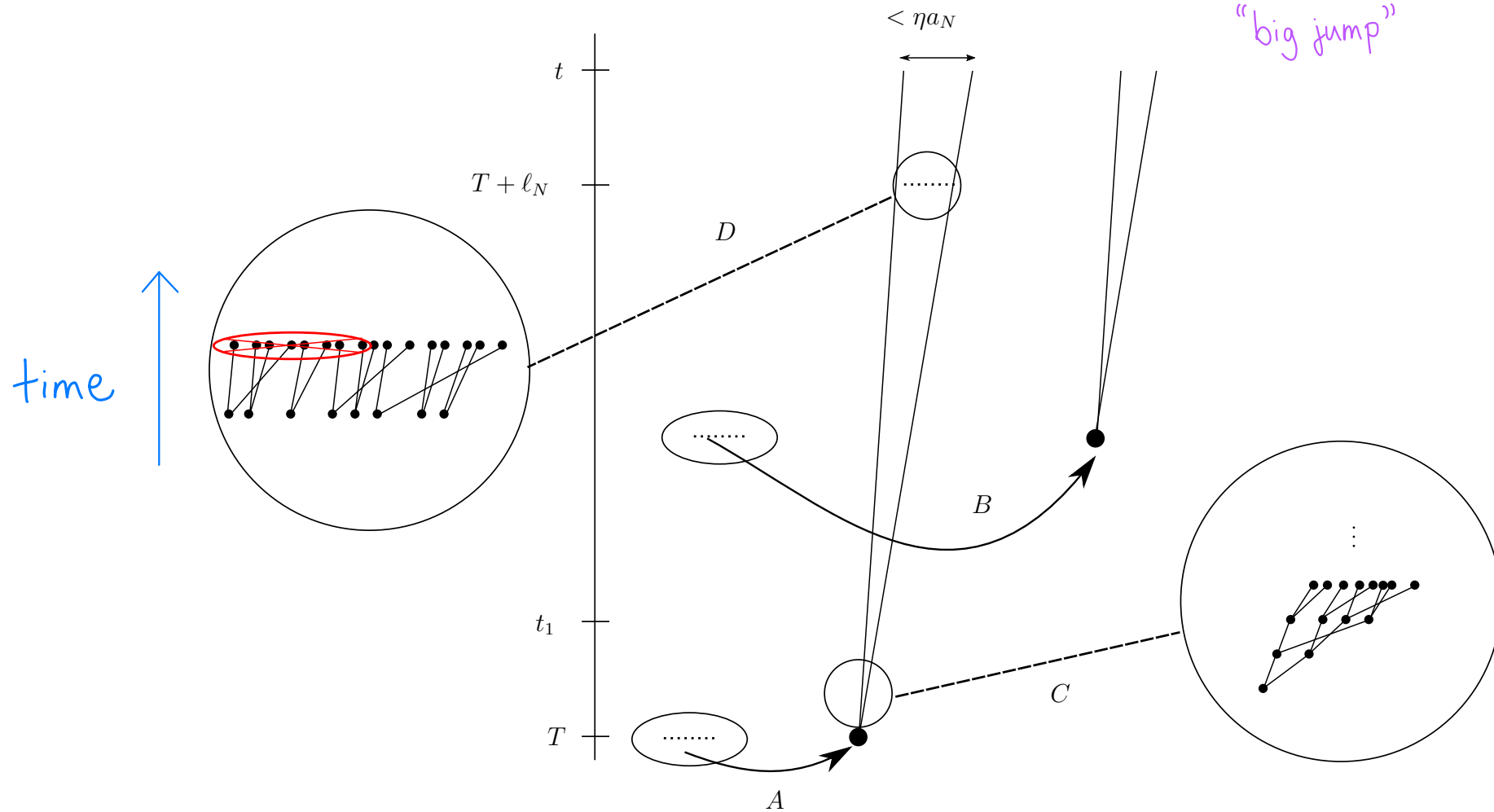
A: A particle makes a big jump at time T and takes the lead (by $\Theta(a_N)$). Its descendants stay in the lead until time t_1 (other particles can't take the lead with a big jump, and can't move far without a big jump).

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B: There are $O(1)$ big jumps in time interval $[t_1, t]$, each with $o(N)$ descendants at time t .

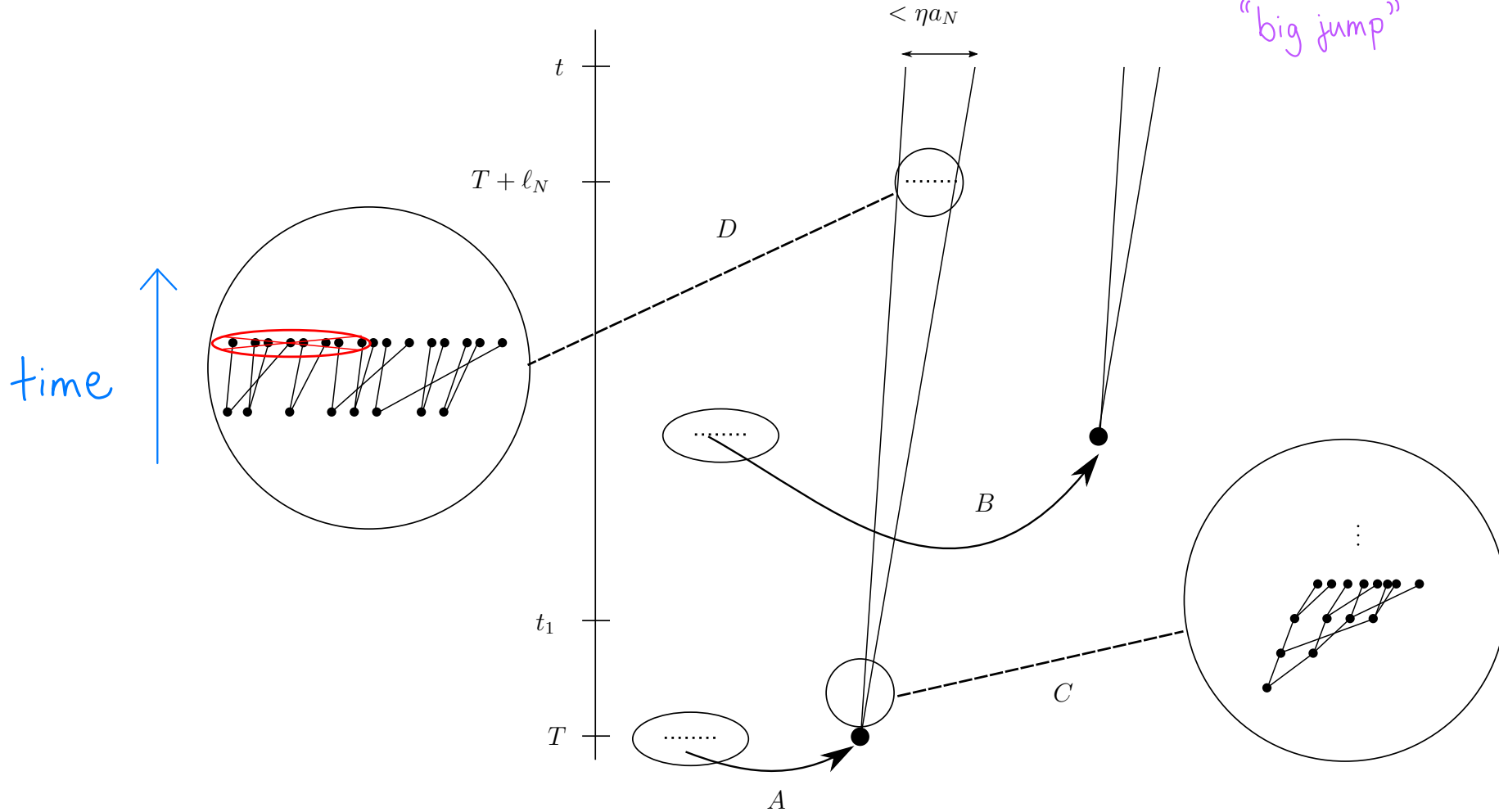
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C: The tribe descended from the time- T leader doubles in size at each timestep until almost time $T + \ell_N$.

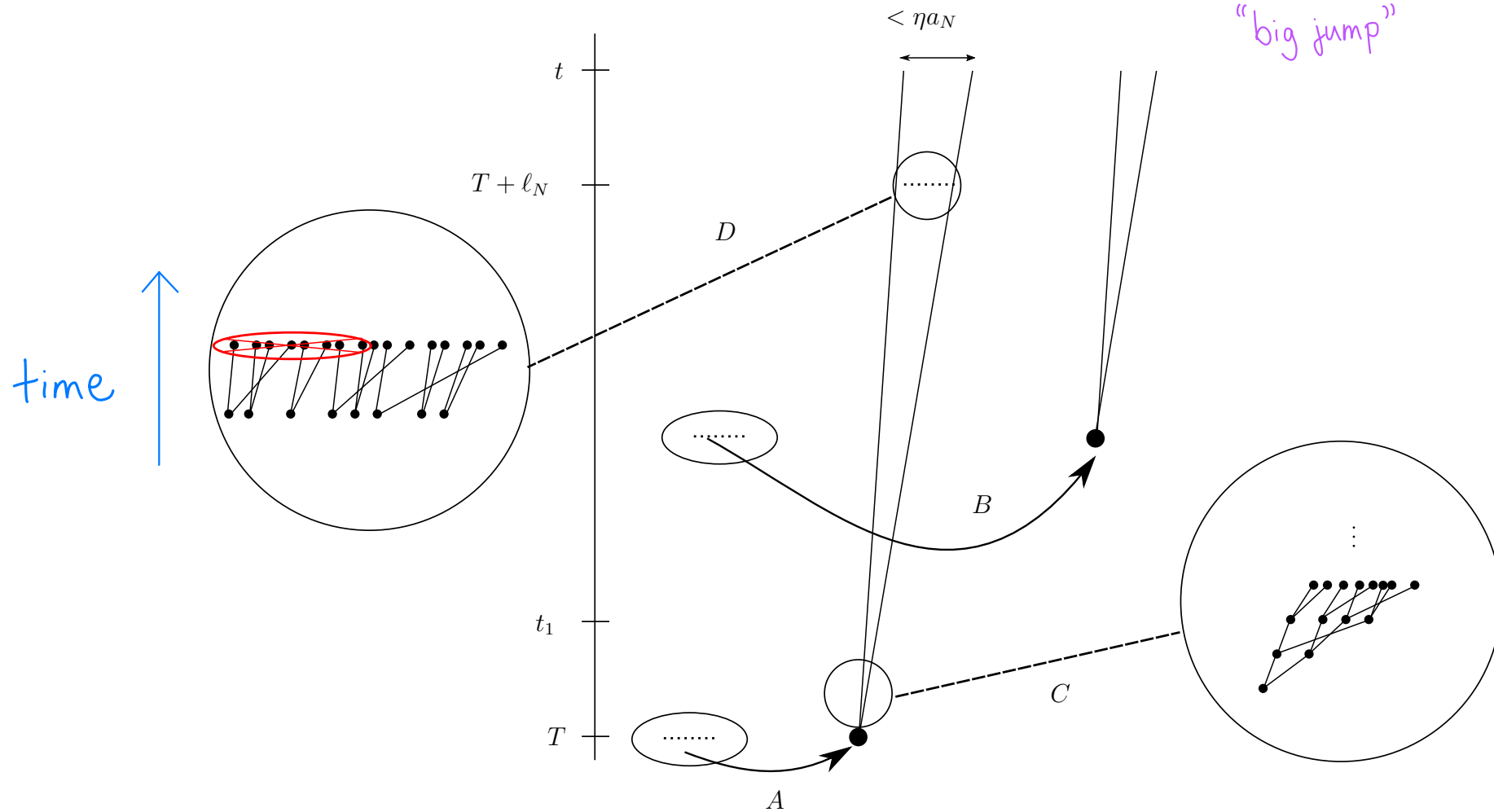
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D: On the time interval $[T + \ell_N, t]$, the time- T leader's tribe has size $N - o(N)$.

N-BRW genealogy

Jump distribution X .

Light-tailed $\mathbb{P}(X > x) \leq e^{-cx}$, $c > 0$

Heavy-tailed $\mathbb{P}(X > x) \sim x^{-\alpha}$, $\alpha > 0$

Time to coalesce

$(\log N)^3$

Coalescent

Bolthausen-Sznitman

$\log N$

Star-shaped

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Work in progress with Z. Talyigás.

N-BRW asymptotic speed

Jump distribution X . $E[X] < \infty$. $v_N := \lim_{n \rightarrow \infty} \frac{X_N^{(N)}(n)}{n}$.

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$$\beta \in (2/3, 1): \quad v_N = \frac{(\log N)^{1/\beta}}{\log_2 N} + \mathbb{E}[X] + \Theta((\log N)^{1-1/\beta})$$

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c.f. Dyzewski, Gantert, Höfelsauer

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up to $\log \log N$ factor

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Simulation by Z. Talyigás.