

§ ... Curvilinear coordinates

So far: Cartesian coordinates

$$\underline{r} = (x, y, z) = (x_1, x_2, x_3)$$

$$\underline{u} = u_1 \underline{e}_1 + u_2 \underline{e}_2 + u_3 \underline{e}_3$$

$\underline{e}_x = \underline{i}$

index: } u_i
notation }


Eqs can be transformed to different coord. systems:

$$\nabla^2 \phi = \frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial y^2} \quad \text{for scalar } \phi(x, y)$$

$$x = r \cos \varphi$$

$$y = r \sin \varphi$$

$$\nabla^2 \phi = \frac{\partial^2 \phi}{\partial r^2} + \frac{1}{r} \frac{\partial \phi}{\partial r} + \frac{1}{r^2} \frac{\partial^2 \phi}{\partial \varphi^2}$$

N. St. eqns contain vectors; 
must transform these &
their derivatives too.

$$\underline{u} = u_1 \underline{e}_1 + u_2 \underline{e}_2 + u_3 \underline{e}_3$$
$$= u_r \underline{e}_r + u_\varphi \underline{e}_\varphi + u_z \underline{e}_z$$

but here the basis vectors
depend on the coords too!

E.g.:

$$\underline{e}_r = \cos\varphi \underline{e}_x + \sin\varphi \underline{e}_y$$

\Rightarrow Differential operators also
act on the basis vectors!

\Rightarrow A MESS!

But we can still use
index notation for
certain eqns:

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$$t_i = \tau_{ij} n_j \quad (i, j) = (r, \phi, z)$$

$$\tau_{ij} = -p \delta_{ij} + 2\mu e_{ij}$$

To illustrate the origin
of some of the terms
in the N-St. eqns in
cyl. polars, consider flow
assoc. with rigid body
rotation.

$$\underline{u} = \underbrace{\omega \phi}_{\Omega r} \underline{e}_\phi \quad ; \quad u_r = u_z = 0$$

\uparrow const.

$$\underline{u} = u_r \underline{e}_r + u_\varphi \underline{e}_\varphi + u_z \underline{e}_z \quad 1$$

$u=0$ $v = \Omega r$ $w=0$

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$$\cancel{\frac{\partial u}{\partial t}} + u \cancel{\frac{\partial u}{\partial r}} + \frac{v \partial u}{r \partial \varphi} + w \cancel{\frac{\partial u}{\partial z}} - \frac{v^2}{r} = -\frac{1}{\rho} \frac{\partial P}{\partial r} + \nu \left[\cancel{\nabla^2 u} - \frac{u}{r^2} - \frac{2}{r^2} \frac{\partial v}{\partial \varphi} \right],$$

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial r} + \frac{v \partial v}{r \partial \varphi} + w \frac{\partial v}{\partial z} + \frac{uv}{r} = -\frac{1}{\rho r} \frac{\partial P}{\partial \varphi} + \nu \left[\nabla^2 v - \frac{v}{r^2} + \frac{2}{r^2} \frac{\partial u}{\partial \varphi} \right],$$

$$\frac{\partial w}{\partial t} + u \frac{\partial w}{\partial r} + \frac{v \partial w}{r \partial \varphi} + w \frac{\partial w}{\partial z} = -\frac{1}{\rho} \frac{\partial P}{\partial z} + \nu \nabla^2 w,$$

$$\text{div } \underline{u} = \frac{1}{r} \frac{\partial}{\partial r} (ru) + \frac{1}{r} \frac{\partial v}{\partial \varphi} + \frac{\partial w}{\partial z} = 0.$$

where

$$\nabla^2 \equiv \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2}{\partial \varphi^2} + \frac{\partial^2}{\partial z^2}.$$

r -comp of mom. eqn gives: 

$$+\frac{v^2}{r} = +\frac{1}{\rho} \frac{dp}{dr}$$

$$\rho \frac{\Omega^2 r^2}{r} = \frac{dp}{dr}$$

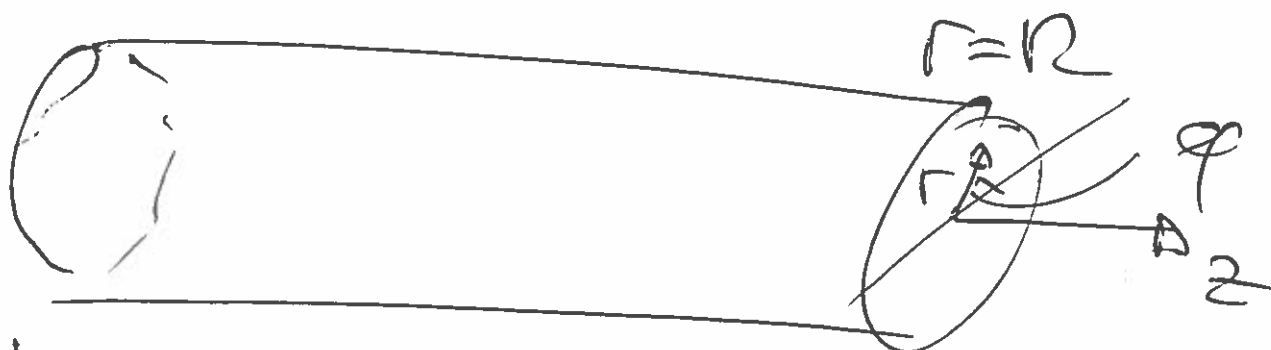
$$\frac{dp}{dr} = \rho \Omega^2 r$$

$$p(r) = p_0 + \frac{1}{2} \rho \Omega^2 r^2$$

This is a centrifugal effect.

Example:

Hagen-Poiseuille flow in pipe



driven by press. drop.

Assume:



$$\underline{u} = u_z \underline{e}_z = \underbrace{u_z(r)}_{\text{(check!)}} \underline{e}_z$$

into r -comp:

$$0 = -\frac{1}{\rho} \frac{dp}{dr} \Rightarrow$$

press does
not depend
on r or ϕ

ϕ -component:

$$0 = -\frac{1}{\rho} \frac{1}{r} \frac{dp}{d\phi}$$

continuity

$$0 = 0 \quad \checkmark$$

z -comp: (check!)

$$0 = -\frac{1}{\rho} \frac{dp}{dz}$$

fcn of z
only.

$$+ r \left(\frac{d^2 \omega}{dr^2} + \frac{1}{r} \frac{d\omega}{dr} \right)$$

fcn of r
only.

As before (parallel flow) $\frac{dp}{dz} = G = \text{const.}$

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$$\frac{G}{\mu} = \underbrace{\frac{\partial^2 \omega}{\partial r^2} + \frac{1}{r} \frac{\partial \omega}{\partial r}}_{\frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial \omega}{\partial r} \right)}$$

$$\frac{G}{\mu} r = \frac{\partial}{\partial r} \left(r \frac{\partial \omega}{\partial r} \right)$$

$$\frac{1}{2} \frac{G}{\mu} r^2 = r \frac{\partial \omega}{\partial r}$$

$$\frac{\partial \omega}{\partial r} = \frac{1}{2} \frac{G}{\mu} r + \frac{A}{r}$$

$$\omega(r) = \frac{1}{4} \frac{G}{\mu} r^2 + A \ln r + B$$

const.
from BC.

No slip: $\omega(r=R) = 0$

& soln is finite @ $r=0$

$$\Rightarrow A = 0$$

$$\omega(r) = \frac{1}{4} \frac{G}{\mu} (r^2 - R^2) \leq 0$$

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