

ONE HOUR THIRTY MINUTES

A list of constants is enclosed.

UNIVERSITY OF MANCHESTER

Particles, Nuclei and Cosmology

27th January 1998, 2.00 p.m. - 3.30 p.m.

Answer TWO questions

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Electronic calculators may be used, provided that they cannot store text.

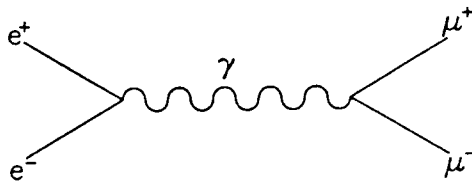
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The numbers indicate the relative weights of the different parts of each question and do NOT represent a marking scheme.

P.T.O.

PC312 January 1998 continued...

1. An electron and a positron annihilate into a virtual photon which then produces a  $\mu^+\mu^-$  pair, as shown below



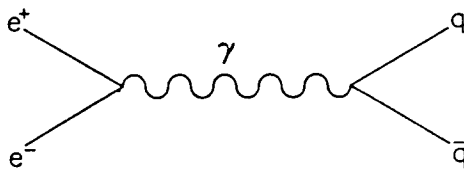
What is the fundamental interaction responsible for this process?

[3 marks]

Explain why the cross-section is proportional to  $\alpha^2$  ( $\alpha$  is the fine structure constant).

[5 marks]

In a similar interaction at the same initial energy a quark-antiquark pair is produced as shown.



Explain why the cross-section for the production of a particular pair of quarks differs from that for  $\mu^+\mu^-$  production by a factor of  $3e_q^2$  where  $e_q$  is the magnitude of the electric charge on the quark (in units of the electron charge).

[6 marks]

If the initial energy is high enough to produce all quark flavours except top, show that if measured at the same initial energy

$$R = \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)} = \frac{11}{3}$$

[7 marks]

Explain briefly how the strong interaction converts the quarks into the experimentally observed hadrons.

[4 marks]

[In units of the electron charge, the electric charges on the  $u, d, s, c, b, t$  quarks are  $+\frac{2}{3}, -\frac{1}{3}, -\frac{1}{3}, +\frac{2}{3}, -\frac{1}{3}, +\frac{2}{3}$ ]

PC312 January 1998 continued...

2. Write down an expression for the mass of a stable nucleus,  $M(Z, A)$  in terms of the masses of the individual proton ( $M_p$ ), and neutron ( $M_n$ ) constituents.

Hence define the term nuclear binding energy.

[4 marks]

Make a sketch of the measured nuclear binding energy in (MeV/nucleon) as a function of mass number  $A$ , showing all the relevant features.

[4 marks]

Describe briefly how the semi-empirical mass formula is evolved from an analogy between the nucleus and a charged liquid drop.

[7 marks]

Show how the main terms in the mass formula can account for the broad dependence of the measured binding energy on  $A$ .

[5 marks]

A more detailed examination of the measured binding energy curve reveals that the nuclei  ${}^4_2\text{He}$ ,  ${}^{16}_8\text{O}$  and  ${}^{208}_{82}\text{Pb}$  have significantly higher binding energies than their near neighbours. Give an explanation for this increased stability.

[5 marks]

3. Discuss at least two pieces of evidence which suggest that we may live in a Universe of finite age.

[8 marks]

Why is it believed that in the early phase of the history of the Universe, a radiation 'fireball' played a dominant role?

[5 marks]

The Universe is imagined to run back in time from the present epoch  $t_0$ , with scale parameter  $R(t_0) = 1$ . At approximately what temperature  $T(t)$ , and scale parameter  $R(t)$  would the temperature begin to rise high enough to provide the 1eV necessary to start the ionisation of some hydrogen atoms? You may assume that  $R(t)$  is inversely proportional to  $T(t)$ . Assuming that the Universe is matter-dominated during the 'run-back' period, how old was the Universe at this temperature? [Take the present age of the Universe to be  $t_0 = 1.5 \times 10^{10}$  yrs.]

[8 marks]

The isotropy and thermal uniformity of the 2.7K cosmic background radiation suggests an equilibrium due to causal connection. How does this give problems to the above model?

[4 marks]