

ONE HOUR THIRTY MINUTES

A list of constants is enclosed.

UNIVERSITY OF MANCHESTER

Introduction to Nuclear and Particle Physics

18th January 2011, 2.00 p.m. - 3.30 p.m.

Answer **ALL** parts of question 1, **ONE** from Section A and **ONE** from Section B

Electronic calculators may be used, provided that they cannot store text.

The numbers are given as a guide to the relative weights of the different parts of each question.

P.T.O.

1. (a) Write down an expression that represents the systematic variation of nuclear radius with mass number A for β -stable nuclei and use it to calculate the radius of an ^{27}Al nucleus.

[5 marks]

- (b) Give an expression relating nuclear mass $m^N(A, Z)$ to atomic mass $m(A, Z)$. Show that the Q value for the β^+ decay process $A \rightarrow B + \beta^+ + \nu_e$ can be written as:

$$Q = m_A - m_B - 2m_e$$

where m_A and m_B are the *atomic* masses of A and B and m_e is the mass of an electron. State clearly any approximations that are made.

[8 marks]

- (c) Explain qualitatively, with the aid of quark flow diagrams, why the charged pions have a lifetime of $\sim 10^{-8}$ s whereas the lifetime of the neutral pion is $\sim 10^{-16}$ s.

[6 marks]

- (d) The particles A and B can be produced in the following strong interaction processes:

$$K^- p \rightarrow \pi^0 A$$

$$K^- p \rightarrow K^+ B.$$

Deduce the charge, baryon, strangeness and charm quantum numbers of the A and B particles, and hence their quark constituents.

The K^- has quark constituents (s, \bar{u}) .

[6 marks]

P.T.O.

SECTION A**Answer either Question A2 or Question A3**

A2. Explain what is meant by *nuclear binding energy*.

[3 marks]

Calculate the binding energy *per nucleon* of a ${}^{62}_{28}\text{Ni}$ nucleus using the following atomic mass excesses: $\Delta({}^{62}\text{Ni}) = -66746.1$, $\Delta({}^1\text{H}) = 7289.0$ and $\Delta(\text{n}) = 8071.3$ keV.

[5 marks]

Draw a diagram indicating the systematic behaviour of the binding energy per nucleon of β -stable nuclei as a function of mass number A . Carefully annotate the axes in a quantitative fashion. Thus explain how energy is generated when: (a) a single heavy nucleus fissions and (b) two light nuclei fuse.

[10 marks]

Sketch a plot showing the variation of nuclear mass with atomic number Z across a sequence of nuclei with the same *even* mass number A . Use your plot to explain why there are often two β -stable isotopes in an even- A isobar.

[5 marks]

There are only four stable odd-odd nuclei: ${}^2\text{H}$, ${}^6\text{Li}$, ${}^{10}\text{B}$ and ${}^{14}\text{N}$. What can be concluded about the effect of pairing in light nuclei?

[2 marks]

P.T.O.

A3. Describe one experimental observation that provides evidence for shell effects in atomic nuclei. [4 marks]

A spin-orbit term with a form proportional to $\mathbf{l} \cdot \mathbf{s}$ is introduced into the single-particle Hamiltonian to reproduce the observed sequence of magic numbers. Show that the expectation of the $\mathbf{l} \cdot \mathbf{s}$ operator can be written as:

$$\langle \mathbf{l} \cdot \mathbf{s} \rangle = \frac{\hbar^2}{2} [j(j+1) - l(l+1) - 3/4].$$

[6 marks]

In a simple three-dimensional potential well, the first three energy levels have quantum numbers $1s$, $1p$ and $1d$ in order of increasing energy. Explain what is meant by this notation.

[4 marks]

Draw an energy level diagram to show the states produced from these levels when the spin-orbit interaction is introduced, labelling the levels with appropriate quantum numbers.

[5 marks]

Use your answers above to explain why the spin and parity of the ground states of ${}^3_2\text{He}$ and ${}^{13}_6\text{C}$ are $1/2^+$ and $1/2^-$ respectively.

[6 marks]

P.T.O.

SECTION B

Answer either Question B4 or Question B5

B4. Briefly describe why a particle such as the neutrino was needed to explain the process of β decay.

[7 marks]

Draw a quark flow diagram for the β decay of a free neutron.

[4 marks]

Why was the experimental discovery of the neutrino so difficult, and how was this eventually achieved?

[5 marks]

Explain the concept of neutrino type.

[6 marks]

What phenomenon associated with neutrinos would lead to the non-conservation of lepton flavour, and what properties would the different types of neutrino need to have for this to occur?

[3 marks]

P.T.O.

B5. Show how the simple 3-flavour quark model can be used to explain the $J=1, L=0$ meson nonet. You do not need to identify the names of particles in this nonet.

[6 marks]

The ρ^- , ρ^0 , ρ^+ , ω^0 and ϕ^0 mesons are the non-strange members of this nonet. Deduce the quark content of each of these mesons.

[7 marks]

Draw quark flow diagrams for the most likely decay channel for the ρ^0 and ϕ^0 mesons.

[8 marks]

Give an estimate of the lifetime and natural width of the ρ^0 .

[4 marks]

The following particle masses may be useful:
 $\rho \approx 780 \text{ MeV}/c^2$, $\omega \approx 780 \text{ MeV}/c^2$, $\phi \approx 1020 \text{ MeV}/c^2$, $\pi \approx 140 \text{ MeV}/c^2$, $K \approx 490 \text{ MeV}/c^2$.

END OF EXAMINATION PAPER

PHYSICAL CONSTANTS AND CONVERSION FACTORS

SYMBOL	DESCRIPTION	NUMERICAL VALUE
c	Velocity of light in vacuum	$299\,792\,458\text{ m s}^{-1}$, exactly
μ_0	Permeability of vacuum	$4\pi \times 10^{-7}\text{ N A}^{-2}$, exactly
ϵ_0	Permittivity of vacuum where $c = \frac{1}{\sqrt{\epsilon_0\mu_0}}$	$8.854 \times 10^{-12}\text{ C}^2\text{ N}^{-1}\text{ m}^{-2}$
h	Planck constant	$6.626 \times 10^{-34}\text{ J s}$
\hbar	$h/2\pi$	$1.055 \times 10^{-34}\text{ J s}$
G	Gravitational constant	$6.674 \times 10^{-11}\text{ m}^3\text{ kg}^{-1}\text{ s}^{-2}$
e	Elementary charge	$1.602 \times 10^{-19}\text{ C}$
eV	Electronvolt	$1.602 \times 10^{-19}\text{ J}$
α	Fine-structure constant, $\frac{e^2}{4\pi\epsilon_0\hbar c}$	$\frac{1}{137.0}$
m_e	Electron mass	$9.109 \times 10^{-31}\text{ kg}$
$m_e c^2$	Electron rest-mass energy	0.511 MeV
μ_B	Bohr magneton, $\frac{e\hbar}{2m_e}$	$9.274 \times 10^{-24}\text{ J T}^{-1}$
R_∞	Rydberg energy $\frac{\alpha^2 m_e c^2}{2}$	13.61 eV
a_0	Bohr radius $\frac{1}{\alpha} \frac{\hbar}{m_e c}$	$0.5292 \times 10^{-10}\text{ m}$
Å	Angstrom	10^{-10} m
m_p	Proton mass	$1.673 \times 10^{-27}\text{ kg}$
$m_p c^2$	Proton rest-mass energy	938.272 MeV
$m_n c^2$	Neutron rest-mass energy	939.565 MeV
μ_N	Nuclear magneton, $\frac{e\hbar}{2m_p}$	$5.051 \times 10^{-27}\text{ J T}^{-1}$
fm	Femtometre or fermi	10^{-15} m
b	Barn	10^{-28} m^2
u	Atomic mass unit, $\frac{1}{12} m(^{12}\text{C atom})$	$1.661 \times 10^{-27}\text{ kg}$
N_A	Avogadro constant, atoms in gram mol	$6.022 \times 10^{23}\text{ mol}^{-1}$
T_t	Triple-point temperature	273.16 K, exactly
k	Boltzmann constant	$1.381 \times 10^{-23}\text{ J K}^{-1}$
R	Molar gas constant, $N_A k$	$8.314\text{ J mol}^{-1}\text{ K}^{-1}$
σ	Stefan-Boltzmann constant, $\frac{\pi^2}{60} \frac{k^4}{\hbar^3 c^2}$	$5.670 \times 10^{-8}\text{ W m}^{-2}\text{ K}^{-4}$
M_E	Mass of Earth	$5.97 \times 10^{24}\text{ kg}$
R_E	Mean radius of Earth	$6.4 \times 10^6\text{ m}$
g	Standard acceleration of gravity	$9.806\,65\text{ m s}^{-2}$, exactly
atm	Standard atmosphere	101 325 Pa, exactly
M_\odot	Solar mass	$1.989 \times 10^{30}\text{ kg}$
R_\odot	Solar radius	$6.96 \times 10^8\text{ m}$
L_\odot	Solar luminosity	$3.84 \times 10^{26}\text{ W}$
T_\odot	Solar effective temperature	$5.8 \times 10^3\text{ K}$
AU	Astronomical unit, mean Earth-Sun distance	$1.496 \times 10^{11}\text{ m}$
pc	Parsec	$3.086 \times 10^{16}\text{ m}$
	Year	$3.156 \times 10^7\text{ s}$