### NUMERICAL STUDY OF A PARTICLE METHOD FOR GRADIENT FLOWS

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ABSTRACT. We study the numerical behaviour of a particle method for gradient flows involving linear and nonlinear diffusion. This method relies on the discretisation of the energy via non-overlapping balls centred at the particles. The resulting scheme preserves the gradient flow structure at the particle level, and enables us to obtain a gradient descent formulation after time discretisation. We give several simulations to illustrate the validity of this method, as well as a detailed study of one-dimensional aggregation-diffusion equations.

#### 1. Introduction

In this work we introduce a new particle method for approximating the solutions to evolution equations of the form

(1.1) 
$$\begin{cases} \rho_t = \nabla \cdot \left[ \rho \nabla \left( H'(\rho(x)) + V(x) + W * \rho(x) \right) \right], & t > 0, x \in \mathbb{R}^d, \\ \rho(0, \cdot) = \rho_0(\cdot), & \end{cases}$$

where  $\rho(t,\cdot) \geq 0$  is the unknown probability measure and  $\rho_0$  is a fixed element of  $\mathcal{P}_2(\mathbb{R}^d)$ , the set of Borel probability measures on  $\mathbb{R}^d$  with bounded second moment. Note that in this paper a probability measure and its density, whenever the latter exists, are denoted by the same symbol. The operator \* denotes the convolution,  $H: [0, \infty) \to \mathbb{R}$  is the density of internal energy,  $V: \mathbb{R}^d \to \mathbb{R}$  is the confinement potential, and  $W: \mathbb{R}^d \to \mathbb{R}$  is the interaction potential. These equations are ubiquitous in many applications, ranging from granular media and porous medium flows to collective behaviour models in mathematical biology and self-assembly, see [6, 8, 12, 31, 32] and the references therein.

Recent advances in the analysis of the equation (1.1) are mainly based on variational schemes using the natural gradient flow structure in the space of probability measures, see e.g. [25, 21, 27, 33, 1, 15]. We define the *continuum energy functional*  $E: \mathcal{P}_2(\mathbb{R}^d) \to \mathbb{R} \cup \{-\infty, +\infty\}$  by

$$E(\rho) = \begin{cases} \int_{\mathbb{R}^d} \left[ H(\rho(x)) + V(x)\rho(x) + \frac{1}{2}W * \rho(x)\rho(x) \right] \mathrm{d}x & \text{if } \rho \in \mathcal{P}_{\mathrm{ac},2}(\mathbb{R}^d), \\ \int_{\mathbb{R}^d} \left[ V(x) + \frac{1}{2}W * \rho(x) \right] \mathrm{d}\rho(x) & \text{if } \rho \not\in \mathcal{P}_{\mathrm{ac},2}(\mathbb{R}^d) \text{ and } H = 0, \\ +\infty & \text{otherwise,} \end{cases}$$

where  $\mathcal{P}_{ac,2}(\mathbb{R}^d)$  is the subset of  $\mathcal{P}_2(\mathbb{R}^d)$  of probability measures which are absolutely continuous with respect to the Lebesgue measure. The functions H,V and W satisfy the following hypotheses.

**Hypothesis 1.** V is a function in  $C^1(\mathbb{R}^d)$  and W is a symmetric, locally integrable function in  $C^1(\mathbb{R}^d \setminus \{0\})$  with W(0) = 0.

**Hypothesis 2.** H is a convex function in  $C^1((0,\infty)) \cap C^0([0,\infty))$  with superlinear growth at infinity and H(0) = 0.

**Hypothesis 3.** The function  $h(\lambda) := \lambda^d H(\lambda^{-d})$  is convex and non-increasing on  $(0, \infty)$ .

Date: 11 November 2015.

The assumption W(0)=0 is made without loss of generality. Indeed, if W(0) is finite, then W can be shifted "up" or "down" to get W(0)=0; if W has a singularity at 0, then setting W(0):=0 does not affect the physical behaviour of a system governed by the potential W. The assumption H(0)=0 and Hypothesis 3 imply that the energy E is displacement convex if, for example, V=W=0, see [24, Section 4] and [33, Theorem 5.15] for a more detailed exposition. Also note that the classical cases with  $H: \rho \mapsto \rho \log \rho$  and  $H: \rho \mapsto \frac{\rho^m}{m-1}$  (m>1) satisfy all the required hypotheses.

The underlying topology on the probability measures in this paper is given by the *quadratic* Wasserstein distance  $d_2(\rho, \mu)$ , which is defined between two measures  $\rho$  and  $\mu$  in  $\mathcal{P}_2(\mathbb{R}^d)$  by

$$d_2(\rho,\mu) = \inf_{\gamma \in \Pi(\rho,\mu)} \left[ \int_{\mathbb{R}^d \times \mathbb{R}^d} |x - y|^2 \, \mathrm{d}\gamma(x,y) \right]^{\frac{1}{2}},$$

where  $\Pi(\rho, \mu)$  is the space of probability measures (also called transport plans) on  $\mathbb{R}^d \times \mathbb{R}^d$  with first marginal  $\rho$  and second marginal  $\mu$ .

Let us fix a final time T > 0. We say that  $\rho: [0,T] \to \mathcal{P}_2(\mathbb{R}^d)$  is a continuum gradient flow solution with initial condition  $\rho_0$  if

(1.2) 
$$\begin{cases} \rho'(t) = -\nabla_{\mathcal{P}_2(\mathbb{R}^d)} E(\rho(t)), \\ \rho(0) = \rho_0, \end{cases}$$

holds in the sense of distributions on  $[0,T] \times \mathbb{R}^d$ , see [1, Theorem 8.3.1]. The operator  $\nabla_{\mathcal{P}_2(\mathbb{R}^d)}$  is the classical quadratic Wasserstein gradient on  $\mathcal{P}_2(\mathbb{R}^d)$ , which takes the following explicit form

$$\nabla_{\mathcal{P}_2(\mathbb{R}^d)} E(\rho) = -\nabla \cdot \left( \rho \nabla \frac{\delta E}{\delta \rho} \right) \quad \text{for all } \rho \in \mathcal{P}_2(\mathbb{R}^d),$$

where  $\frac{\delta E}{\delta \rho} = H'(\rho) + V + W * \rho$  is the first variation density of E at point  $\rho$ . As a by-product of the general theory developed for instance in [1], gradient flow solutions to (1.2) are weak solutions to (1.1). Note that if theoretical issues as the existence and uniqueness of solutions to the continuum gradient flow (1.2) are of interest, appropriate additional assumptions must be imposed on E and  $\rho_0$ , see [1, Theorem 11.1.14 and Section 11.1.3].

We propose below to approximate solutions to the continuum gradient flow (1.2) by finite atomic probability measures represented by finite numbers of particles. The basic idea is to restrict the continuum gradient flow to the discrete setting of atomic measures, by performing the steepest descent of a suitable approximation of the energy E defined on finite numbers of Dirac masses, as a discrete analogue to the JKO scheme for the Fokker-Planck equation proposed by Jordan, Kinderlehrer and Otto [21]. The theoretical underpinning of this method is studied in the companion paper [17], where the convergence of the discrete gradient flow to the continuum one is proved in the framework proposed by Serfaty in [30, 29], in the special case of V = W = 0 in one dimension and under additional appropriate hypotheses on H. The goal of the present paper is to give numerical evidences of such convergence and to motivate possible extensions of the theoretical result in [17] to nonzero confinement and interaction potentials (even with possible singularities), as well as to higher dimensions.

Let us mention that other numerical methods have been developed to conserve particular properties of solutions of the continuity equation (1.1). In [7, 12] the authors developed finite volume methods preserving the decay of energy at the semi-discrete level, along with other important properties like non-negativity and mass conservation. Particle methods for these equations without the diffusion term are known to be convergent under suitable assumptions on the potentials V and W since these results are connected to the question of the mean-field limit and rigorous derivation of the equations from particle trajectories, see [13] for instance. In the case of diffusion equations, particle methods based on suitable regularisations of the flux of the continuity equation (1.1) have

been proposed in [28, 18, 22, 23]. Our steepest descent method is purely variational and is based on regularising the internal part of the energy E by substituting particles by non-overlapping blobs, as we discuss next. Finally, let us mention that the numerical approximation of the JKO variational scheme has already been tackled by different methods using pseudo-inverse distributions in one dimension [19, 8, 10], diffeomorphisms [16], or solving for the optimal map in a JKO step [5]. Our method avoids these computationally intensive procedures by the approximation of the energy in the discrete setting.

In Section 2 we give a summary of the method and the derivation of the discrete gradient flow in a more general setting than the one considered in [17]. Section 3 is dedicated to the derivation of the numerical scheme used to approximate the continuum gradient flow (1.2), that is an explicit JKO scheme, as well as to a numerical validation study of the scheme via diffusion equations. Finally, in Section 4 we give the numerical results for various one-dimensional aggregation-diffusion equations.

## 2. Particle method and discrete gradient flows

In this method, the underlying probability measure is characterised by the particles' positions  $(x_1,\ldots,x_N)\in(\mathbb{R}^d)^N=\mathbb{R}^{Nd}$  and the associated weights (or masses)  $w:=(w_1,\cdots,w_N)\in\mathbb{R}^N$ , where N is the total number of particles considered. Throughout this paper, the positions  $(x_1,\ldots,x_N)$  are evolving in time but the weights w are fixed and such that  $w_i>0$  and  $\sum_{i=1}^N w_i=1$ . Also, we denote by  $\mathbb{R}^{Nd}_w$  the space of particles with weights w, that is,  $\boldsymbol{x}:=(x_1,\ldots,x_N)\in\mathbb{R}^{Nd}_w$  means that each particle  $x_i$  is in  $\mathbb{R}^d$  and is associated with the weight  $w_i$ . Note the boldface font when referring to elements of  $\mathbb{R}^{Nd}_w$ .

Remark 2.1. As an important convention, in the rest of the paper, whenever particles  $\boldsymbol{x} \in \mathbb{R}_w^{Nd}$  are considered, they are assumed to be distinct, i.e.,  $x_i \neq x_j$  if  $i \neq j$ . Moreover, in one dimension, the particles are assumed to be sorted increasingly, i.e.,  $x_{i+1} > x_i$  for all  $i \in \{1, \ldots, N-1\}$ .

2.1. Discrete gradient flow. Consider N particles  $x := (x_1, \dots, x_N) \in \mathbb{R}_w^{Nd}$  with fixed weights w. The most natural representation of the underlying probability measure is the empirical measure

(2.1) 
$$\mu_N = \sum_{i=1}^N w_i \delta_{x_i},$$

which belongs to the space

$$\mathcal{A}_{N,w}(\mathbb{R}^d) := \left\{ \mu \in \mathcal{P}_2(\mathbb{R}^d) \mid \exists \, \boldsymbol{x} \in \mathbb{R}_w^{Nd}, \mu = \sum_{i=1}^N w_i \delta_{x_i} \right\}.$$

**Definition 2.2** (Discrete energy). We define the discrete energy  $E_N : \mathcal{A}_{N,w}(\mathbb{R}^d) \to \mathbb{R}$  by

(2.2) 
$$E_N(\mu_N) = \sum_{i=1}^N |B_i| H\left(\frac{w_i}{|B_i|}\right) + \sum_{i=1}^N w_i V(x_i) + \frac{1}{2} \sum_{i=1}^N \sum_{\substack{j=1 \ j \neq i}}^N w_i w_j W(x_i - x_j),$$

where  $B_i$  is the open ball of centre  $x_i$  and radius  $\frac{1}{2}\min_{j\neq i}|x_i-x_j|$ , and  $|B_i|$  is the volume of  $B_i$ .

Note that  $E_N$  is finite over the whole  $\mathcal{A}_{N,w}(\mathbb{R}^d)$  since, by the Hypotheses 1 and 2, H,V and W are pointwise finite. The essence of this discrete approximation  $E_N$  on  $\mathcal{A}_{N,w}(\mathbb{R}^d)$  of the continuum energy E on  $\mathcal{P}_2(\mathbb{R}^d)$  lies in the treatment of the internal part  $\int_{\mathbb{R}^d} H(\rho(x)) dx$  of the energy E, which becomes infinity on point masses; the point mass of each particle is spread to circumvent this problem. To this end, consider

(2.3) 
$$\rho_N = \sum_{i=1}^N w_i \frac{\chi_{B_i}}{|B_i|},$$

where  $\chi_{B_i}$  is the characteristic function on the ball  $B_i$ . Clearly  $\rho_N$  is in  $\mathcal{P}_{ac,2}(\mathbb{R}^d)$ , and thus the internal part of the energy is well-defined for  $\rho_N$ . Obviously, the representation  $\rho_N$  does not involve overlapping of balls, but contains "gaps" between balls whose sizes are expected to decrease as the number of particles increases. Then, using (2.3) for the internal part of the energy and (2.1) for the confinement and interaction parts,  $E_N$  defined in (2.2) is exactly

$$\int_{\mathbb{R}^d} H(\rho_N(x)) dx + \int_{\mathbb{R}^d} V(x) d\mu_N(x) + \int_{\mathbb{R}^d \times \mathbb{R}^d} W(x-y) d\mu_N(y) d\mu_N(x).$$

Here the diagonal terms in the interaction potential vanish since W(0) = 0 by Hypothesis 1.

Since the expression above depends essentially on  $\boldsymbol{x} \in \mathbb{R}^{Nd}_w$ , we can define the discrete energy equivalently as a function of  $\boldsymbol{x} \in \mathbb{R}^{Nd}_w$  instead of  $\mu_N \in \mathcal{A}_{N,w}(\mathbb{R}^d)$ :

(2.4) 
$$\widetilde{E}_N(\boldsymbol{x}) := E_N(\mu_N)$$
 for all  $\mu_N \in \mathcal{A}_{N,w}(\mathbb{R}^d)$  with particles  $\boldsymbol{x} \in \mathbb{R}^{Nd}_w$ .

**Definition 2.3** (Subdifferential in Hilbert spaces). Let X be a Hilbert space and  $\phi: X \to \mathbb{R}$ . We define the *subdifferential*  $\partial \phi: X \to 2^X$  of  $\phi$ , for all  $x \in X$ , by the subset

$$\partial \phi(x) = \left\{ z \in X \mid \liminf_{y \to x} \frac{\phi(y) - \phi(x) - \langle z, y - x \rangle_X}{|y - x|_X} \ge 0 \right\}.$$

If  $\partial \phi(x)$  is not empty, then  $\partial^0 \phi(x)$  is defined to be the unique element in  $\partial \phi(x)$  with minimal norm.

Now, we say that  $x:[0,T]\to\mathbb{R}^{Nd}_w$  is a discrete gradient flow solution with initial condition  $x^0\in\mathbb{R}^{Nd}$  if the following differential inclusion

(2.5) 
$$w \cdot \mathbf{x}'(t) \in -\partial \widetilde{E}_N(\mathbf{x}(t))$$
 for almost all  $t \in (0, T]$ ,

is satisfied, and  $\mathbf{x}(0) = \mathbf{x}^0$ . Here  $\mathbf{x}'(t)$  is the velocity of the curve  $\mathbf{x}(t)$  and  $w \cdot \mathbf{x}'$  is the element-wise product  $(w_1 x_1', \dots, w_N x_N')$ . This formulation is not a standard differential inclusion because of the appearance of the weights. To cope with this, we introduce the following inner product on  $\mathbb{R}_w^{Nd}$ .

**Definition 2.4** (Weighted inner product on  $\mathbb{R}_w^{Nd}$ ). For all  $x, y \in \mathbb{R}_w^{Nd}$  we define the weighted inner product between x and y as

$$\langle \boldsymbol{x}, \boldsymbol{y} \rangle_w = \sum_{i=1}^N w_i \langle x_i, y_i \rangle_{\mathbb{R}^d}.$$

From now on, the Euclidean space  $\mathbb{R}_w^{Nd}$  is endowed with this inner product. This definition clearly induces the following weighted norm on  $\mathbb{R}_w^{Nd}$ ,

(2.6) 
$$|\mathbf{x}|_w := \sqrt{\langle \mathbf{x}, \mathbf{x} \rangle_w} = \sqrt{\sum_{i=1}^N w_i |x_i|_{\mathbb{R}^d}^2} for all \mathbf{x} \in \mathbb{R}_w^{Nd}.$$

It also induces a slightly more general subdifferential structure: for any functional  $\phi \colon \mathbb{R}_w^{Nd} \to \mathbb{R}$ ,

$$\partial_w \phi(m{x}) := \left\{ m{z} \in \mathbb{R}_w^{Nd} \mid w \cdot m{z} \in \partial \phi(m{x}) 
ight\} \quad ext{for all } m{x} \in \mathbb{R}_w^{Nd},$$

where again  $w \cdot z$  is the element-wise product between  $w \in \mathbb{R}^N$  and  $z \in \mathbb{R}^{Nd}_w$ ; and we can then define the element  $\partial_w^0 \phi(x)$  with minimal norm accordingly. The discrete gradient flow inclusion (2.5) can now be rewritten more naturally as

(2.7) 
$$\begin{cases} \boldsymbol{x}'(t) \in -\partial_w \widetilde{E}_N(\boldsymbol{x}(t)) & \text{for almost all } t \in (0, T], \\ \boldsymbol{x}(0) = \boldsymbol{x}^{\boldsymbol{0}}. \end{cases}$$

This is the discrete gradient flow structure that is used from now on.

Remark 2.5. In [17] the authors showed that in one dimension with V = W = 0 the discrete energy  $E_N$   $\Gamma$ -converges (in the  $d_2$ -topology, see Definition A.3) to the continuum one E, which they used to show that the discrete gradient flow converges to the continuum one in the sense proposed by Serfaty in [30] if adequate additional assumptions are imposed on H.

2.2. p-approximated discrete gradient flow. The gradient flow (2.7) is written as a differential inclusion instead of an ordinary differential equation, primarily because the radius  $\frac{1}{2} \min_{j \neq i} |x_i - x_j|$  of the balls  $B_i$  in (2.3) is not a smooth function of the positions in the definition of the energy  $\tilde{E}_N$ . As a result, the energy  $\tilde{E}_N$  is not a smooth function of the positions x, which especially happens when the radius of the ball  $B_i$  is determined by more than one particle.

If  $\widetilde{E}_N$  is lower semi-continuous and convex, and  $\boldsymbol{x^0} \in \{\boldsymbol{y} \in \mathbb{R}_w^{Nd} \mid \partial_w \widetilde{E}_N(\boldsymbol{y}) \neq \emptyset\}$ , it is known by [4, Theorem 1 of Section 3.2] that the discrete gradient flow (2.7) is well-posed and the equation  $\boldsymbol{x'}(t) = -\partial_w^0 \widetilde{E}_N(\boldsymbol{x}(t))$  is satisfied for almost every  $t \in (0,T]$ . Note that, in our case, the lower semi-continuity of  $\widetilde{E}_N$  is trivial to check by the assumptions on H, whereas the convexity is proved in Proposition A.6 for convex potentials V and W in one dimension. Although the element of minimal norm  $\partial_w^0 \widetilde{E}_N(\boldsymbol{x}(t))$  can be computed in some situations as in [17] (for the special case with internal energy only and in one dimension), the detailed procedure is usually involved and is not convenient for numerical implementation. In theory, one can approximate  $\widetilde{E}_N$  by a differentiable function usually characterised by a parmeter, where the minimal norm element of the subdifferential is recovered as the limit of the gradient of the approximation as the parameter goes to infinity. A common choice of such an approximation is

$$Y_N^p(oldsymbol{x}) := \inf_{oldsymbol{y} \in \mathbb{R}_w^{Nd}} \left( \widetilde{E}_N(oldsymbol{y}) + rac{p}{2} |oldsymbol{x} - oldsymbol{y}|_w^2 
ight),$$

depending on the parameter p > 0. Then  $Y_N^p$  is differentiable almost everywhere in  $\mathbb{R}_w^{Nd}$ ,  $\nabla Y_N^p$  is called the *Yosida approximation* of  $\partial_w \widetilde{E}_N$ , and

$$abla Y_N^p(oldsymbol{x}) \xrightarrow[p o \infty]{} \partial_w^0 \widetilde{E}_N(oldsymbol{x}) \quad ext{for all } oldsymbol{x} \in \mathbb{R}_w^{Nd},$$

see [4, Theorem 2 of Section 3.1 and Theorem 4 of Section 3.4]. However, the Yosida approximation requires another optimisation, and is therefore not very well adapted to the computation of  $\partial_w^0 \widetilde{E}_N(\boldsymbol{x})$ . An alternative is to approximate the radius of the balls  $B_i$  in the definition of  $\widetilde{E}_N$  by a smooth function.

**Definition 2.6** (p-approximation of the minimum function). For p > 0 and integer  $s \ge 2$ , the function  $\min_p : (0, \infty)^s \to (0, \infty)$  is defined as

$$\min_{p}(x_1, \dots, x_s) = \begin{cases} \left(\frac{1}{s} \sum_{i=1}^s x_i^{-p}\right)^{-\frac{1}{p}} & \text{if } x_1, \dots, x_s \in (0, \infty), \\ 0 & \text{otherwise.} \end{cases}$$

We call  $\min_p$  the *p-approximation* of the minimum function  $\min(x_1, \dots, x_s)$ .

The following proposition, whose proof is straightforward and left to the reader, justifies the use of the p-approximation of the minimum function.

**Proposition 2.7.** Let  $s \geq 2$ . The following statements hold.

- (1) Let p > 0. Then  $\min_{p \in C^{\infty}((0, \infty)^s)}$ .
- (2)  $\min_{n}(x) \to \min(x)$  as  $p \to \infty$  for all  $x \in (0, \infty)^{s}$ .
- (3)  $p \mapsto \min_{p}(x)$  is non-increasing on  $(0, \infty)$  for all  $x \in (0, \infty)^{s}$ .
- (4) Let p > 0. Then  $\min_{p}(x) \ge \min(x)$  for all  $x \in (0, \infty)^s$ , with equality if  $x_i = x_j$  for all i, j.
- (5) Let  $p \ge 1$ . Then  $\min_{p}(x) \le s^{1/p} \min(x)$  for all  $x \in (0, \infty)^s$ .

Now the discrete energy is further approximated using a smoothed radius as below.

**Definition 2.8** (p-approximated discrete energy). Let p > 0. The p-approximated discrete energy  $E_N^p \colon \mathcal{A}_{N,w}(\mathbb{R}^d) \to \mathbb{R}$  is defined by

(2.8) 
$$E_N^p(\mu_N) = \sum_{i=1}^N |B_i^p| H\left(\frac{w_i}{|B_i^p|}\right) + \sum_{i=1}^N w_i V(x_i) + \frac{1}{2} \sum_{i=1}^N \sum_{\substack{j=1\\j \neq i}}^N w_i w_j W(x_i - x_j),$$

for all  $\mu_N \in \mathcal{A}_{N,w}(\mathbb{R}^d)$  with particles  $\boldsymbol{x} \in \mathbb{R}^{Nd}_w$ . Here  $|B_i^p|$  is the volume of the ball with the new radius

$$\frac{1}{2}\min_{p}\{|x_i - x_j| \mid j \neq i\} = \frac{1}{2} \left( \frac{1}{N-1} \sum_{\substack{j=1\\j \neq i}}^{N} |x_i - x_j|^{-p} \right)^{-\frac{1}{p}},$$

where the minimisation is performed on the neighbours of  $x_i$  (at most two neighbours if d = 1).

As in (2.4), we can define the p-approximated discrete energy on  $\mathbb{R}^{Nd}_w$  rather than on  $\mathcal{A}_{N,w}(\mathbb{R}^d)$ :

(2.9) 
$$\widetilde{E}_N^p(\boldsymbol{x}) := E_N^p(\mu_N) \text{ for all } \mu_N \in \mathcal{A}_{N,w}(\mathbb{R}^d) \text{ with particles } \boldsymbol{x} \in \mathbb{R}_w^{Nd}.$$

We say that  $x:[0,T]\to\mathbb{R}^{Nd}_w$  is a *p-approximated discrete gradient flow solution* with initial condition  $x^0\in\mathbb{R}^{Nd}_w$  if

(2.10) 
$$\begin{cases} \boldsymbol{x}'(t) = -\nabla_w \widetilde{E}_N^p(\boldsymbol{x}(t)) & \text{for almost all } t \in (0, T], \\ \boldsymbol{x}(0) = \boldsymbol{x}^0, \end{cases}$$

where the weighted gradient  $\nabla_w$  is naturally defined on  $\mathbb{R}^{Nd}_w$  by

$$\nabla_w := \left(\frac{1}{w_1} \frac{\partial}{\partial x_1}, \dots, \frac{1}{w_N} \frac{\partial}{\partial x_N}\right).$$

Remark 2.9. Whenever d=1 and V and W are convex, we know by Proposition A.6 that  $\widetilde{E}_N^p$  is convex, and therefore the p-approximated discrete gradient flow (2.10) is well-posed on [0,T].

Remark 2.10. The interest of the p-approximation of the discrete gradient flow only lies in the numerical simplicity of coding a gradient descent on an ODE system such as (2.10) rather than on a differential inclusion such as (2.7). The p-approximation is indeed not needed for the theoretical proof given in [17] of the convergence of the discrete gradient flow to the continuum one, since, as already mentioned, in that specific case the minimal norm element of  $\partial \widetilde{E}_N$  is explicitly computable.

Remark 2.11. Approximations of the minimum function other than the one given in Definition 2.6 are possible. One example is

$$\min_{p}(x_1, \dots, x_s) = \begin{cases} -\frac{1}{p} \log \left( \frac{1}{s} \sum_{i=1}^s e^{-px_i} \right) & \text{if } x_1, \dots, x_s \in (0, \infty), \\ 0 & \text{otherwise.} \end{cases}$$

The properties of  $\min_p$  given in Proposition 2.7 are not enough to say that the p-approximated discrete gradient flow (2.10) converges to the discrete one (2.7) as  $p \to \infty$  in some sense. This still needs to be checked if we want to justify the numerical use of the p-approximated gradient flow. In the next section we prove this is the case in dimension one since we can there exploit the convexity of the discrete energies.

#### 2.3. One-dimensional case.

**Definition 2.12** (Inter-particle distance). For any particles  $x \in \mathbb{R}_w^N$  we denote the *inter-particle distance* by the positive quantity (eventually  $+\infty$  by convention)

$$\Delta x_i := x_i - x_{i-1} \quad \text{for } i \in \{1, \dots, N+1\},$$

with the convention  $x_0 = -\infty$  and  $x_{N+1} = +\infty$ . We also write, for p > 0,

$$r_i = \min(\Delta x_i, \Delta x_{i+1})$$
 and  $r_i^p = \min_p(\Delta x_i, \Delta x_{i+1})$  for all  $i \in \{1, \dots, N\}$ .

Note that  $r_i = |B_i|$  and  $r_i^p = |B_i^p|$ .

Since particles are sorted increasingly by convention, the p-approximated discrete energy (2.8) in dimension one can be defined in the simpler form

(2.11) 
$$E_N^p(\mu_N) = \sum_{i=1}^N r_i^p H\left(\frac{w_i}{r_i^p}\right) + \sum_{i=1}^N w_i V(x_i) + \frac{1}{2} \sum_{i=1}^N \sum_{\substack{j=1\\j \neq i}}^N w_i w_j W(x_i - x_j),$$

for all  $\mu_N \in \mathcal{A}_{N,w}(\mathbb{R})$  with particles  $\boldsymbol{x} \in \mathbb{R}_w^{Nd}$ . The equivalent formulation of the energy  $\widetilde{E}_N^p$  on the particles is defined in dimension one accordingly, see (2.9).

In order to check that the p-approximated discrete gradient flow defined above indeed approximates the discrete one, we need a few elements of maximal monotone operator theory (see [3, 4, 9] for a detailed overview of the theory). These notions and results being quite abstract, they are only given in Appendix A. There we show how to use this theory to prove the convergence of the p-approximated discrete gradient flow to the discrete one in a precise sense, in the case when V and W are convex; and we therefore justify the numerical use of the p-approximated gradient flow (2.10) (at least in the case when V and W are convex and d = 1).

#### 3. Numerical scheme and validation

Here a general variational formulation is applied to the discrete setting described above, leading to our numerical scheme giving the particles' positions at discrete time steps (see [1, 2, 21] for a Fokker-Planck motivation and its generalisation to curves in probability spaces). For the sake of generality the derivation of the scheme is partly done for the discrete gradient flow, rather than the p-approximated one. The simulations given later were, however, performed in the p-approximated setting only. Note also that, in this section, some subindices N are dropped for clarity.

3.1. The scheme. Take  $(t_n)_{n=0}^M \subset [0,T]$  a subdivision of the time interval [0,T] and a timestep size  $\Delta t = \Delta_n t$  (eventually adaptive) such that  $t_n = t_0 + \sum_{i=0}^{n-1} \Delta_i t$  for all  $n \in \{1,\ldots,M\}$ . Suppose that, for some  $n \in \{0,\ldots,M-1\}$ , we know  $\mu^n \in \mathcal{A}_{N,w}(\mathbb{R}^d)$ , the empirical measure of an approximated discrete gradient flow solution at time  $t_n$ . Then we want to get an approximated discrete gradient flow solution at the time  $t_{n+1}$ , that is  $\mu^{n+1}$ . To this end we use the JKO scheme, i.e., we choose  $\mu^{n+1}$  as

(3.1) 
$$\mu^{n+1} := \underset{\mu \in \mathcal{A}_{N,w}(\mathbb{R}^d)}{\operatorname{argmin}} \left\{ \frac{1}{2\Delta t} d_2^2(\mu^n, \mu) + E_N(\mu) \right\},$$

where we recall that  $d_2$  is the quadratic Wasserstein distance. In the following any object with subor superscript n or n + 1 is associated with the time step n or n + 1, respectively.

From now on in this paper, apart from a two-dimensional test in Section 3.4.5, the setting is one-dimensional.

3.1.1. Computation of the Wasserstein distance. Let us compute the Wasserstein distance in (3.1). For each  $n \in \{0, ..., M-1\}$  and approximation  $\mu^n$  and  $\mu \in \mathcal{A}_{N,w}(\mathbb{R})$ , we directly get

(3.2) 
$$d_2^2(\mu^n, \mu) = \sum_{i=1}^N w_i (x_i^n - x_i)^2,$$

where  $\boldsymbol{x}^n = (x_1^n, \dots, x_N^n)$  and  $\boldsymbol{x} = (x_1, \dots, x_N)$  are the particles of  $\mu^n$  and  $\mu$ , respectively, which leads to the scheme

(3.3) 
$$\mu^{n+1} := \underset{\mu \in \mathcal{A}_{N,w}(\mathbb{R})}{\operatorname{argmin}} \left( \sum_{i=1}^{N} w_i \frac{(x_i^n - x_i)^2}{2\Delta t} + E_N(\mu) \right).$$

Clearly the scheme (3.3) on the empirical measures can be equivalently rewritten on the particles:

(3.4) 
$$w \cdot \boldsymbol{x}^{n+1} := \operatorname*{argmin}_{\boldsymbol{x} \in \mathbb{R}_w^N} \left( \sum_{i=1}^N w_i \frac{(x_i^n - x_i)^2}{2\Delta t} + \widetilde{E}_N(\boldsymbol{x}) \right),$$

where  $w \cdot x^{n+1}$  is the element-wise multiplication between vectors w and  $x^{n+1}$ . The minimisation problem (3.4) is characterised by

(3.5) 
$$0 \in \partial_w \left( \sum_{i=1}^N w_i \frac{(x_i^n - x_i^{n+1})^2}{2\Delta t} + \widetilde{E}_N(\boldsymbol{x}^{n+1}) \right).$$

When  $\widetilde{E}_N$  is convex, (3.5) is precisely the backward Euler scheme of the differential inclusion (2.7).

3.1.2. Minimisation and final form of the scheme. Let p > 0. We get back to the p-approximated setting, i.e., we consider (3.4) (and equivalently (3.3)) with  $\widetilde{E}_N^p$  instead of  $\widetilde{E}_N$  (an equivalently  $E_N^p$  instead of  $E_N$ ). We want to minimise, over the whole set  $\mathbb{R}_w^N$ , the functional in the argmin operator in (3.4) (and (3.3)) to find our approximation  $\boldsymbol{x}^{n+1}$  (and  $\boldsymbol{\mu}^{n+1}$ ) at time step n+1. First note that, for each  $n \in \{0, \ldots, M-1\}$ , we can write the discrete energy  $\widetilde{E}_N^p$  as

$$\widetilde{E}_{N}^{p}(\boldsymbol{x}^{n+1}) = \sum_{i=1}^{N} w_{i} E_{i}^{n+1},$$

where

$$E_i^{n+1} := r_i^{n+1} H\left(\frac{w_i}{r_i^{n+1}}\right) + w_i V(x_i^{n+1}) + \frac{1}{2} \sum_{\substack{j=1\\j\neq i}}^N w_j W(x_i^{n+1} - x_j^{n+1}),$$

where  $r_i^{n+1} := r_i^{p,n+1}$ , see Definition 2.12. For any  $n \in \{0,\ldots,M-1\}$ , (3.5) becomes

(3.6) 
$$x_i^{n+1} = x_i^n - \frac{\Delta t}{w_i} \sum_{j=1}^N w_j \frac{\partial E_j^{n+1}}{\partial x_i^{n+1}},$$

where the uninspiring computation of the derivatives in the sum-terms is left to the reader. The scheme (3.6) is the implicit Euler scheme of the ODE (2.10), or equivalently, of the ODE

$$w_i \frac{\mathrm{d}x_i}{\mathrm{d}t} = -\frac{\partial \widetilde{E}_N^p}{\partial x_i}(\boldsymbol{x}), \quad i = 1, 2, \dots, N,$$

and it coincides with the JKO scheme for  $\widetilde{E}_N^p$ , as in (3.1). Since the implicit scheme (3.6) is difficult and costly to solve, its explicit Euler version was used in the numerical examples of this paper:

(3.7) 
$$x_i^{n+1} = x_i^n - \frac{\Delta t}{w_i} \sum_{j=1}^N w_j \frac{\partial E_j^n}{\partial x_i^n}.$$

3.2. **Initialisation of the scheme.** The approximation of the initial condition is important. Below, we give two different ways of approximating the initial profile, but let us first introduce the notion of pseudo-inverse.

**Definition 3.1** (Pseudo-inverse). Let  $F: \mathbb{R} \to [0,1]$  be a non-decreasing function. The *pseudo-inverse* of F is the non-decreasing function defined by  $\Phi: [0,1] \to \mathbb{R} \cup \{-\infty, +\infty\}$  and

$$\Phi(\varepsilon) = \inf\{x \in \mathbb{R} \mid F(x) > \varepsilon\} \text{ for all } \varepsilon \in [0, 1].$$

3.2.1. Initially equally weighted particles. If we want to approximate the initial profile  $\rho_0 \in \mathcal{P}_2(\mathbb{R})$  properly with equally weighted particles, we need to start with unequally spaced particles. Consider the pseudo-inverse  $\Phi_0$  of the cumulative distribution  $x \mapsto \rho_0((-\infty, x])$  of the initial profile. Suppose  $w_i = 1/N$  for all  $i \in \{1, ..., N\}$ . Then choose the initial particles  $\boldsymbol{x^0} := (x_1^0, ..., x_N^0) \in \mathbb{R}_w^N$  as

$$x_i^0 = \Phi_0\left(\frac{2i-1}{2N}\right)$$
 for all  $1 \le i \le N$ .

3.2.2. Initially equally spaced particles. If now we want to approximate the initial profile  $\rho_0 \in \mathcal{P}_2(\mathbb{R})$  properly with equally spaced particles, we need to a assign different weight to each particle. Consider  $F_0$ , the cumulative distribution of the initial profile, and  $x^0 \in \mathbb{R}^N_w$ , some chosen equally spaced initial particles. Let us define

$$x_{i\pm\frac{1}{2}}^0 = \tfrac{1}{2} \left( x_i^0 + x_{i\pm1}^0 \right) \quad \text{for all } 1 \leq i \leq N.$$

Then we choose the weights w as

(3.8) 
$$\begin{cases} w_1 = F_0\left(x_{1+\frac{1}{2}}^0\right), \\ w_i = F_0\left(x_{i+\frac{1}{2}}^0\right) - F_0\left(x_{i-\frac{1}{2}}^0\right) & \text{for all } 1 < i < N, \\ w_N = 1 - F_0\left(x_{N-\frac{1}{2}}^0\right). \end{cases}$$

The weight  $w_i$  as given in (3.8) for each 1 < i < N is the mass that  $\rho_0$  "assigns" to the interval  $[x_{i-\frac{1}{2}}, x_{i+\frac{1}{2}}]$ . Note that the choice we have on the initial positions of the particles  $\boldsymbol{x}^0$  is not completely free of constraints: they must be chosen inside the support of  $\rho_0$  or the resulting weights are null. Finally, since the initial particles are chosen to be equally spaced, only the first and last particles' locations are needed to determine the locations of all the others; for the following numerical simulations we give this information under the form  $I_{\text{init}} = [x_1, x_N]$ .

Remark 3.2. There are two main advantages in the second initialisation approach. The first one is that it allows a better approximation when there is a strong variation in the density; indeed one has the freedom to place initial particles in lower populated regions, allowing a better approximation of the initial profile. The second one is that there are no gaps between the initial balls of centres  $x_i^0$  and diameters  $\min_p(\Delta x_i^0, \Delta x_{i+1}^0) = \min(\Delta x_i^0, \Delta x_{i+1}^0)$ , see Proposition 2.7(4), allowing again a better approximation of the initial profile.

- 3.3. Computation of the error. The natural error to compute for our scheme (3.7) is the quadratic Wassertein error.
- 3.3.1. Error with respect to an exact solution. If we want to get the error between the numerical and exact solutions we proceed as follows. Let  $\rho \in \mathcal{P}_2(\mathbb{R})$  be a continuum gradient flow solution at the final time T. Also write  $\boldsymbol{x} \in \mathbb{R}^N_w$ , the approximation of the p-approximated discrete gradient flow solution obtained with the JKO scheme (3.7) at the final time step M with associated empirical measure  $\mu := \mu^M$ . Then we define the quadratic Wasserstein error by

$$e_{d_2} = d_2(\rho, \mu).$$

To compute this, one can use the one-dimensional pseudo-inverse definition of the quadratic Wasserstein distance. Let us write F and G the cumulative distributions of  $\rho$  and  $\mu$ , respectively, and  $\Phi$  and  $\Psi$  their respective pseudo-inverses. Then

$$e_{d_2} = \left(\int_0^1 \left(\Phi(\varepsilon) - \Psi(\varepsilon)\right)^2 d\varepsilon\right)^{\frac{1}{2}}.$$

We have, for all  $i \in \{1, ..., N\}$ ,

$$\Psi(\varepsilon) = x_i \quad \text{if} \quad \varepsilon \in [\Omega_{i-1}, \Omega_i),$$

with  $\Omega_i := \sum_{j=1}^i w_j$  and the convention  $\Omega_0 = 0$ . Therefore

(3.9) 
$$e_{d_2} = \left(\sum_{i=1}^{N} \int_{\Omega_{i-1}}^{\Omega_i} (x_i - \Phi(\varepsilon))^2 d\varepsilon\right)^{\frac{1}{2}},$$

The error can therefore be easily determined if the inverse of the cumulative distribution of the exact solution at the final time T is known.

3.3.2. Error with respect to a discrete steady state. If we know that the considered gradient flow has a steady state and we are interested in the stabilisation behaviour of the scheme (3.7) we proceed as follows. Let  $\boldsymbol{x}^* := (x_1^*, \dots, x_N^*) \in \mathbb{R}_w^N$  be a discrete steady state (obtained by running a simulation for a "large" time) with associated empirical measure  $\mu^*$ , and let  $\boldsymbol{x} \in \mathbb{R}_w^N$  be the approximation of the p-approximated discrete gradient flow solution obtained with the JKO scheme (3.7) at some final time step M with empirical measure  $\mu := \mu^M$ . Then we define the error by

(3.10) 
$$e_{d_2}^* = d_2(\mu^*, \mu) = \left(\sum_{i=1}^N w_i (x_i^* - x_i)^2\right)^{\frac{1}{2}},$$

which is nothing but the *Euclidean error* on the particle space  $\mathbb{R}^N_w$ , see (2.6).

- 3.4. **Numerical validation: Diffusions.** Before giving any simulation results, let us summarise practical implementation aspects of the scheme.
- As already justified in Section 3.1, all the following simulations were obtained by implementing the explicit scheme (3.7).
- The parameter p needed for the iteration of such scheme was chosen to be 10 for every simulation. This choice is only justified by the fact that it gave sensible results. Note that taking p "large" is a bad idea. Indeed, if p increases, then the p-approximated discrete gradient flow "approaches" the discrete one, for which the energy  $\widetilde{E}_N$  is not differentiable whenever two neighbouring interparticle distances are equal. This may cause numerical instabilities as we would expect the inter-particle distances at the centre of a symmetric profile (like a Gaussian) to be indeed equal; this is what we observed for large p under the form of particle oscillations.
- All one-dimensional simulations were initialised in the way described in Section 3.2.2. In each case, the initial continuum profile and the interval  $I_{\text{init}}$  are explicitly given.
- The choice of the time-step size is explicitly given for each simulation.
- All the solution profiles given in the figures for the one-dimensional simulations were drawn by linking linearly the centres of each constant piece of the function  $\rho_N$  defined in (2.3). Note that this is only one way of representating the discrete solution, and other choices are also possible.

In this section we validate our scheme (3.7) by showing test simulations run on diffusion equations: the heat and porous medium equations.

3.4.1. The heat equation. The (linear) heat equation  $\rho_t = \Delta \rho$  is a continuum gradient flow (1.2) for  $H(\rho) = \rho \log \rho$  and V = W = 0. The evolution of the solution starting from the initial data

(3.11) 
$$\rho_0^{\text{heat}}(x) = \frac{1}{\sqrt{4\pi t_0}} e^{-\frac{x^2}{4t_0}} \quad \text{with } t_0 = 0.25$$

is shown in Figure 1a. The exact solution is given, for all t > 0, by

$$\rho(t,x) = \frac{1}{\sqrt{4\pi(t+t_0)}} e^{-\frac{x^2}{4(t+t_0)}} \text{ for all } x \in \mathbb{R},$$

whose cumulative distribution F and its pseudo-inverse  $\Phi$  are

$$\begin{cases} F(t,x) = \frac{1}{2} \left( 1 + \operatorname{erf}\left(\frac{x}{\sqrt{4(t+t_0)}}\right) \right) & \text{for all } x \in \mathbb{R}, \\ \Phi(t,\varepsilon) = \sqrt{4(t+t_0)} \operatorname{erf}^{-1}(2\varepsilon - 1) & \text{for all } \varepsilon \in [0,1], \end{cases}$$

where erf is the error function. The error at the final time T can then be found using (3.9) and some quadrature form to approximate the integrals therein, see Figure 1b. In Figure 1a we chose  $I_{\text{init}} = [-2.5, 2.5]$ , whereas in Figure 1b we chose  $I_{\text{init}} = [-4, 4]$ , see Section 3.2.2.

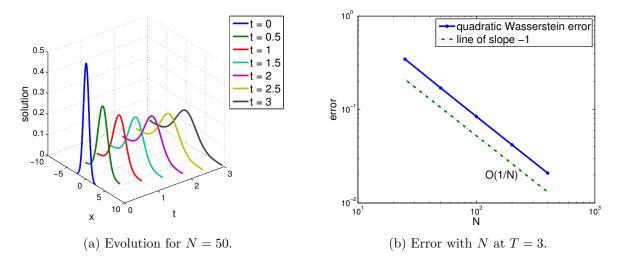


FIGURE 1. The heat equation with  $\Delta t = 10^{-5}$ .

From Figure 1b we can see that the quadratic Wasserstein error with respect to the exact solution  $\rho$  evolves linearly with the number of particles on a log-log plot. From this plot, it looks fair to say that the error of our scheme (3.7) is  $e_{d_2} = \mathcal{O}\left(\frac{1}{N}\right)$ .

3.4.2. The porous medium equation. The porous medium equation  $\rho_t = \Delta \rho^m$  is a continuum gradient flow (1.2) for  $H(\rho) = \frac{\rho^m}{m-1}$ , m > 1, and V = W = 0. The evolution of the solution starting from the initial data

(3.12) 
$$\rho_0^{\text{por}}(x) = \frac{1}{t_0^{\alpha}} \psi\left(\frac{x}{t_0^{\alpha}}\right) \text{ with } t_0 = 0.25$$

is shown in Figure 2a. Here  $\alpha = \frac{1}{m+1}$ ,  $\psi \colon \xi \mapsto (K - \kappa \xi^2)_+^{1/(m-1)}$ , where the subscript + stands for the positive part,  $\kappa = \frac{m-1}{2m(m+1)}$  and K determined by the conservation of mass. Indeed, since the total conserved mass is one, then the constant K can be expressed as

$$K = \left[\Gamma\left(\frac{1}{m-1} + \frac{3}{2}\right)\sqrt{\kappa} / \left(\Gamma\left(\frac{m}{m-1}\right)\Gamma\left(\frac{1}{2}\right)\right)\right]^{\frac{2(m-1)}{m+1}},$$

where  $\Gamma$  is the Gamma-function. Then one can verify that, for all t > 0,

$$\rho(t,x) = \frac{1}{(t+t_0)^{\alpha}} \psi\left(\frac{x}{(t+t_0)^{\alpha}}\right) \quad \text{for all } x \in \mathbb{R},$$

is a solution, see [32, Section 4.4]. Moreover, for  $x \in \left(0, t^{\alpha} \sqrt{\frac{K}{\kappa}}\right]$ , the cumulative distribution is

$$F(t,x) = \frac{1}{2} + \int_0^x \frac{1}{t^{\alpha}} \left( K - \frac{\kappa x^2}{t^{2\alpha}} \right)_+^{\frac{1}{m-1}} dx = \frac{1}{2} + \frac{1}{2} B\left(1; \frac{1}{2}, \frac{m}{m-1}\right)^{-1} B\left(\frac{\kappa}{K} \frac{x^2}{t^{2\alpha}}; \frac{1}{2}, \frac{m}{m-1}\right),$$

where B is the incomplete Beta-function:

$$B(x; a, b) = \int_0^x z^{a-1} (1-t)^{b-1} dz$$
 for all  $x \ge 0$  and  $a, b > 0$ .

Similarly, for  $x \in \left(-t^{\alpha}\sqrt{\frac{K}{\kappa}}, 0\right]$ ,

$$F(t,x) = \frac{1}{2} - \frac{1}{2}B\left(1; \frac{1}{2}, \frac{m}{m-1}\right)^{-1}B\left(\frac{\kappa}{K}\frac{x^2}{t^{2\alpha}}; \frac{1}{2}, \frac{m}{m-1}\right).$$

Also, noticing that

$$K^{\frac{1}{m-1} + \frac{1}{2}} \kappa^{-\frac{1}{2}} = \left( B\left(1; \frac{1}{2}, \frac{m}{m-1}\right) \right)^{-1},$$

we get, using the normalised incomplete Beta-function  $I(x; a, b) = \frac{B(x; a, b)}{B(1; a, b)}$ , that the pseudo-inverse of F is

$$\Phi(t,\varepsilon) = \begin{cases} -t^{\alpha} \sqrt{\frac{K}{\kappa}} \sqrt{I^{-1} \left(1 - 2\varepsilon; \frac{1}{2}, \frac{m}{m-1}\right)} & \text{if } 0 \le \varepsilon < \frac{1}{2}, \\ t^{\alpha} \sqrt{\frac{K}{\kappa}} \sqrt{I^{-1} \left(2\varepsilon - 1; \frac{1}{2}, \frac{m}{m-1}\right)} & \text{if } \frac{1}{2} \le \varepsilon < 1, \\ +\infty & \text{if } \varepsilon = 1. \end{cases}$$

The error can then be found using (3.9) and approximating the integrals therein, see Figure 2b. In Figure 2 we chose  $I_{\text{init}} = \left[ -t_0^{\alpha} \sqrt{K/\kappa}, t_0^{\alpha} \sqrt{K/\kappa} \right]$ , see Section 3.2.2.

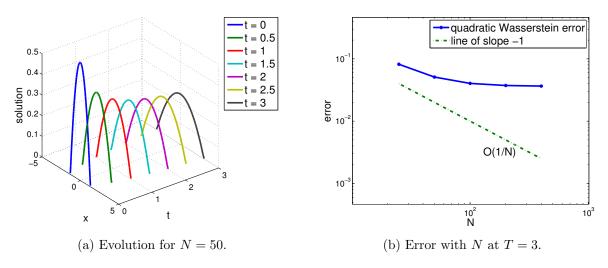


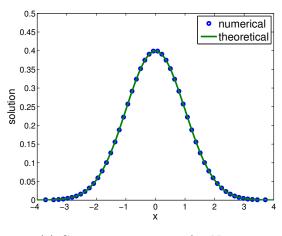
FIGURE 2. The porous medium equation with m=2 and  $\Delta t=\frac{0.1}{N^2}$ .

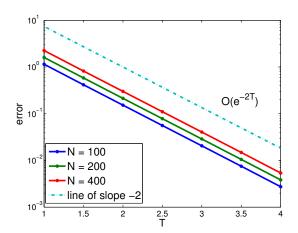
From Figure 2b we can see that for a certain range of N, the quadratic Wasserstein error with respect to the exact solution  $\rho$  evolves fairly linearly with the number of particles on a log-log plot. From this plot, as for the one in Figure 1b for the heat equation, it looks fair to say again that the error of our scheme (3.7) is  $e_{d_2} = \mathcal{O}\left(\frac{1}{N}\right)$ . The deviation from a straight line appearing in Figure 2b for large numbers of particles (and which could occur as well for the heat equation in Figure 1b for even larger numbers) can be explained as follows. The way we chose to initialise our set of particles (see Section 3.2.2) is very likely far from the optimal way of approximating the initial profile, especially as N increases. This initial error was then "dragged" as time increased and seems to have a greater effect on the final error of nonlinear equations as the porous medium equation than on linear ones as the heat equation. The term "optimal" here means "best possible as to minimise the error with respect to  $d_2$ ". Optimising the error between a combination of point masses and a continuum profile is a classical problem (see quantisation theory, for example [20]) which we do not deal with in the present paper. Another reason for this discrepancy in the error for the porous medium equation comes from the fact that the initial profile is compactly supported, and therefore, since the initial particles must be chosen inside this support, see Section 3.2.2, they get very packed from the start as N increases.

3.4.3. The linear Fokker-Planck equation. Let us consider the heat equation with quadratic confinement potential, i.e., the heat equation with  $V(x) = \frac{x^2}{2}$ . In this case, regardless of the initial condition, there is a steady state:

$$\rho_{\infty}(x) := \frac{1}{\sqrt{2\pi}} e^{-\frac{x^2}{2}} \text{ for all } x \in \mathbb{R}.$$

Figure 3b was obtained from the continuum initial profile  $\rho_0^{\text{heat}}$ , see (3.11), with  $I_{\text{init}} = [-2.5, 2.5]$  for Figure 3a and  $I_{\text{init}} = [-4, 4]$  for Figure 3b.





(a) Comparison at T = 4 for N = 50.

(b) Error with final time T for different N's with respect to the discrete steady state.

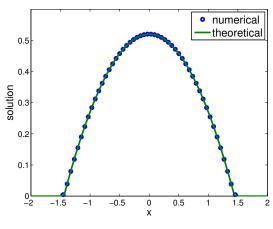
FIGURE 3. The linear Fokker-Planck equation with  $\Delta t = 10^{-5}$  – Stabilisation in time of the scheme (rate of convergence to the discrete steady state).

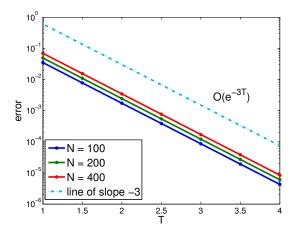
In Figure 3b, we can see that the stabilisation of the scheme towards the discrete steady state, which was found by running the scheme up to T=6, is linear on a semi-log plot with a slope very close to -2, which seems not to depend on the number of particles N. We can therefore write the error as  $e_{d_2}^* = \mathcal{O}\left(e^{-2T}\right)$ , see (3.10).

3.4.4. The nonlinear Fokker-Planck equation. Let us consider the porous medium equation with quadratic confinement potential, i.e., the porous medium equation with  $V(x) = \frac{x^2}{2}$ . In this case, regardless of the initial profile and since the total mass is one, the steady state is

$$\rho_{\infty}(x) := A(R^2 - x^2)_{+}^{\frac{1}{m-1}} \quad \text{for all } x \in \mathbb{R},$$

where  $A=(\frac{m-1}{2m})^{1/(m-1)}$  and  $R=(AB(1;\frac{1}{2},\frac{m}{m-1}))^{(1-m)/(m+1)}$ . Figure 4b was obtained from the continuum initial profile  $\rho_0^{\rm por}$ , see (3.12), with  $I_{\rm init}=\left[-t_0^{\alpha}\sqrt{K/\kappa},t_0^{\alpha}\sqrt{K/\kappa}\right]$ .





- (a) Comparison at T = 4 for N = 50.
- (b) Error with the final time T for different N's with respect to the discrete steady state.

FIGURE 4. The nonlinear Fokker-Planck equation with m=2 and  $\Delta t = \frac{0.1}{N^2}$  – Stabilisation in time of the scheme (rate of convergence to the discrete steady state).

As already noted from Figure 3b, for the heat equation the numerical error can be written  $e_{d_2}^* = \mathcal{O}\left(e^{-2T}\right)$ , as it is actually expected from the theory; for the porous medium equation the theory says that we should expect  $e_{d_2}^* = \mathcal{O}\left(e^{-(m+1)T}\right)$ , see [14] and references therein, which is nicely confirmed numerically by Figure 4b with m=2.

3.4.5. A two-dimensional test: the heat equation. It is straightforward to generalise the scheme (3.7), derived in Section 3.1, to higher dimensions whenever the expression (3.2) for the Wasserstein distance can be used (which is the case when the approximation  $\mu^{n+1}$  at time step n+1 is sufficiently close to  $\mu^n$ ). When the time step  $\Delta t$  is small, (3.2) can still be considered as correct up to a "small" error, since the possible splitting of mass occurring in higher-dimensional optimal transport is then negligible.

Here we test the scheme for the (linear) heat equation in two dimensions. The initial continuum density is  $\rho_0(x) = \frac{1}{4\pi t_0} e^{-|x|^2/(4t_0)}$  with  $t_0 = 0.125$ , and the particle positions at T = 1 are shown in Figure 5b. The data were initialised by fixing the particles on a regular grid as in Figure 5a, with weights calculated as the integral of the continuum density  $\rho_0$  on the Voronoi cells generated from the particles (see for instance [20] for a detailed account on Voronoi cells).

Certain statistics along the time evolution, like the second moments  $\sum_{i=1}^{N} w_i |x_i|^2$  and the entropy  $\sum_{i=1}^{N} w_i \log \frac{w_i}{|B_i|}$ , seem to be very accurate, as shown in Figure 6. However, representing the numerical solution in this two-dimensional test is a delicate issue since gaps between discretisation balls are significant; this is an issue in itself, which we do not deal with here.

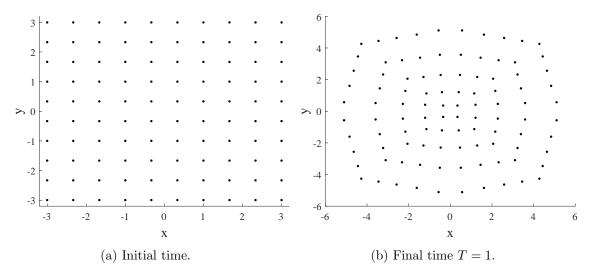


FIGURE 5. The particles' positions for the two-dimensional heat equation for N = 100, with  $\Delta t = 10^{-4}$ .

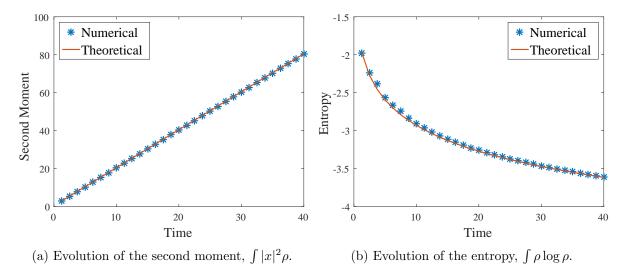


Figure 6. Accuracy for the two-dimensional heat equation, with  $\Delta t = 10^{-4}$ .

# 4. Aggregation-Diffusion Equations

4.1. The modified Keller-Segel equation. We start by considering a modified one-dimensional Keller-Segel equation, that is the continuum gradient flow (1.2) with  $H(\rho) = \rho \log \rho$  and  $W(x) = 2\chi \log |x|$  (and W(0) := 0), where  $\chi > 0$  is a parameter quantifying the attraction (see [11, 8] for well-posedness and qualitative properties, and [10] for the approximation of this equation by a different particle method).

This model shows a critical behaviour depending only on the chemosensitivity strength  $\chi$  as the classical Keller-Segel model in two dimensions, that is there is a dichotomy of blow-up in finite time or global existence which is only determined by  $\chi$  being larger or less than 1, see [11, 8]. In case  $\chi < 1$  solutions spread in time behaving like self-similar solutions. To get the leading order profile given by the self-similar solution, a time-space scaling is done for  $\chi < 1$ , which is equivalent to impose a quadratic confinement potential on particles, see [8]. The long-time behaviour in this subcritical case is given by the profile of the self-similar solution.

4.1.1. Theoretical properties. We show that our particle approximation keeps approximately the criticality of the original Keller-Segel model at the discrete level. We show that there exist two positive constants  $\chi_1$  and  $\chi_2(N)$  such that the following holds: if an appropriate confinement potential V is considered, the discrete and the p-approximated discrete gradient flows (2.7) and (2.10) for the modified Keller-Segel equation have steady states for  $\chi < \chi_1$ ; while if V = 0, the p-approximated discrete gradient flow shows finite-time blow-up for  $\chi > \chi_2(N)$ . Quite surprisingly,  $\chi_1$  happens to be exactly the critical parameter at the continuum level, i.e.,  $\chi_1 = 1$ , and  $\chi_2(N)$  tends to 1 as  $N \to \infty$  and does not depend on p, but only on N and the chosen weights.

By the term "blow-up" at the discrete level we mean the event of two, or more, particles colliding. Also note that in the following the term "maximal time of existence" indicates either the first time when two or more particles of a solution collide, i.e., the first blow-up time, or the first time when the norm of a solution equals  $+\infty$ .

First, let us prove that the discrete and p-approximated discrete confined Keller-Segel equations show no collisions of particles if  $\chi < \chi_1$ .

**Proposition 4.1.** Consider the discrete gradient flow corresponding to the confined Keller-Segel equation with V symmetric and such that there exist  $\beta > (w_1 + w_N)^{-1}$  and  $R_{\infty} > 0$  with  $V(x) > \beta \log |x|$  for all  $|x| > R_{\infty}$ . Suppose there exists a solution  $\mathbf{x}$  to such gradient flow, emanating from an initial condition  $\mathbf{x}^0 \in \mathbb{R}^N_w$ , up to some maximal time of existence, say  $T^* > 0$ . If

$$\chi < \chi_1 := 1$$
,

then no particles of x can collide in  $[0,T^*)$ ; furthermore, the minimal inter-particle distance is uniformly bounded from below in time by a positive constant.

*Proof.* Let us assume that  $x_1(t) = -x_N(t) < 0$  for all  $t \in [0, T^*)$  without loss of generality, by simply translating the space frame appropriately. The energy  $\widetilde{E}_N$  is a Lyapunov functional, i.e.,

(4.1) 
$$\widetilde{E}_N(\boldsymbol{x}(t)) \leq \widetilde{E}_N(\boldsymbol{x}^0) := E_0 < +\infty \text{ for all } t \in [0, T^*),$$

see [4, Theorem 1 of Section 3.4]. Fix  $t \in [0, T^*)$  and get, by (2.2),

(4.2) 
$$\widetilde{E}_{N}(\boldsymbol{x}(t)) = \sum_{i=1}^{N} w_{i} \log w_{i} - \sum_{i=1}^{N} w_{i} \log r_{i}(t) + \sum_{i=1}^{N} w_{i} V(x_{i}(t)) + \sum_{i=1}^{N} \sum_{\substack{j=1\\ i \neq i}}^{N} w_{i} w_{j} \log |x_{i}(t) - x_{j}(t)|.$$

Writing  $\log_{-} x := \log x$  if 0 < x < 1 and  $\log_{-} x := 0$  if  $x \ge 1$ , and using that  $\log$  is increasing,

$$\sum_{i=1}^{N} \sum_{\substack{j=1\\j\neq i}}^{N} w_i w_j \log |x_i(t) - x_j(t)| \ge \sum_{i=1}^{N} \sum_{\substack{j=1\\j\neq i}}^{N} w_i w_j \log \min_{\substack{k \in \{1,\dots,N\}\\k\neq i}} |x_i(t) - x_k(t)|$$

$$\ge \sum_{i=1}^{N} w_i (1 - w_i) \log_{-} r_i(t) \ge \sum_{i=1}^{N} w_i \log_{-} r_i(t),$$

where  $r_i(t) = \min(\Delta x_i(t), \Delta x_{i+1}(t))$ . Hence, writing  $\log_+ x := \log x$  if  $x \ge 1$  and  $\log_+ x := 0$  if  $0 \le x < 1$ , and by (4.2),

$$\widetilde{E}_N(\boldsymbol{x}(t)) \ge \sum_{i=1}^N w_i \log w_i + \sum_{i=1}^N w_i V(x_i(t)) + (\chi - 1) \sum_{i=1}^N w_i \log_- r_i(t) - \sum_{i=1}^N w_i \log_+ r_i(t).$$

Define the function

$$f(x) = (w_1 + w_N)V(x) - \log_+(2x)$$
 for all  $x \in [0, \infty)$ .

We then get

$$\widetilde{E}_{N}(\boldsymbol{x}(t)) \geq \sum_{i=1}^{N} w_{i} \log w_{i} + \sum_{i=2}^{N-1} w_{i} V(x_{i}(t)) + w_{1} V(x_{1}(t)) + w_{N} V(x_{N}(t))$$

$$+ (\chi - 1) \sum_{i=1}^{N} w_{i} \log_{-} r_{i}(t) - \log_{+} (x_{N}(t) - x_{1}(t))$$

$$= \sum_{i=1}^{N} w_{i} \log w_{i} + \sum_{i=2}^{N-1} w_{i} V(x_{i}(t)) + (\chi - 1) \sum_{i=1}^{N} w_{i} \log_{-} r_{i}(t) + f(x_{N}(t)),$$

using that  $-\log_+ r_i(t) \ge -\log(x_N(t) - x_1(t))$  for all  $i \in \{1, \dots, N\}$ , the symmetry of V and that  $x_1(t) = -x_N(t)$ . Now, by the growth assumption at infinity on V we see that  $f(x) \to +\infty$  as  $|x| \to +\infty$ , and  $f(0) = (w_1 + w_N)V(0) > -\infty$ . Therefore, by continuity of f, there exists a constant  $C_1 \in \mathbb{R}$  such that  $f(x) \ge C_1$  for all  $x \in [0, \infty)$ . Then, since from the assumptions V is bounded from below, there exists a constant  $C_2 \in \mathbb{R}$ , independent of t, such that

(4.3) 
$$E_0 \ge \widetilde{E}_N(\boldsymbol{x}(t)) \ge (\chi - 1) \sum_{i=1}^N w_i \log_- r_i(t) + C_1 + C_2,$$

using (4.1). We see that if  $\chi < \chi_1 = 1$ , the minimal inter-particle distances  $r_i(t)$  cannot get arbitrarily small, or the energy  $\widetilde{E}_N(\boldsymbol{x}(t))$  gets larger than its initial value  $E_0$ , which violates the fact that the system is a gradient flow. Hence result.

Remark 4.2. The proof above is given only for the discrete case; however, note that it can be easily adapted to the p-approximated one, if  $p \ge 1$ , by Proposition 2.7(5) with s = 2, and without the need to change the constant  $\chi_1$ .

We can now show the global existence in time and the existence of steady states for the discrete and p-approximated discrete confined Keller-Segel equations.

**Proposition 4.3.** Consider the discrete gradient flow corresponding to the confined Keller-Segel equation with V satisfying the same hypotheses as in Proposition 4.1. If  $\chi < \chi_1$ , then any solution to this gradient flow, if it exists, exists globally in time and the gradient flow has a steady state.

Proof. Let  $\chi < \chi_1 = 1$ . Suppose there exists such a solution  $\boldsymbol{x}$ , emanating from an initial condition  $\boldsymbol{x^0} \in \mathbb{R}^N_w$ , defined on  $[0,T^*)$ , see Proposition 4.1, and assume that  $x_N(t) - x_1(t) \to +\infty$  as  $t \to T^*$ . As in the proof of Proposition 4.1, we can suppose that  $x_1(t) = -x_N(t) < 0$  for all  $t \in [0,T^*)$ , yielding  $x_1(t) \to -\infty$  and  $x_N(t) \to +\infty$  as  $t \to T^*$ . For all  $t \in [0,T^*)$ , the proof of Proposition 4.1 implies that there exists a t-independent constant  $C_2 \in \mathbb{R}$  such that

$$E_0 \ge \widetilde{E}_N(\boldsymbol{x}(t)) \ge (\chi - 1) \sum_{i=1}^N w_i \log_- r_i(t) + f(x_N(t)) + C_2 \ge f(x_N(t)) + C_2,$$

since  $(\chi - 1) \sum_{i=1}^{N} w_i \log_{-} r_i(t) \geq 0$ . In addition, since  $x_N(t) \to +\infty$  as  $t \to T^*$ , we know that  $f(x_N(t)) \to +\infty$ , which implies that the inequality above is violated for a time t close enough to  $T^*$ . Therefore  $x_N(t) - x_1(t)$  cannot diverge as  $t \to T^*$ , and thus the set  $\{x_1(t), \ldots, x_N(t)\}$  has to be uniformly bounded in time, i.e., there exists some constant  $\ell > 0$ , independent of time, such that  $\{x_1(t), \ldots, x_N(t)\} \subset B(0, \ell)$ . This, together with the no-collision result in Proposition 4.1, shows that the maximal time of existence of the discrete gradient flow solution is  $T^* = \infty$ .

Finally, the functional  $E_N$  is lower semi-continous and bounded from below on any sublevel set of  $\tilde{E}_N$  which are compact due to (4.3). Indeed, the previous argument shows that  $x_i \in B(0, \ell)$  for

all  $\boldsymbol{x} \in \mathbb{R}_w^N$  such that  $E_0 \geq \widetilde{E}_N(\boldsymbol{x})$ , and thus  $r_i \leq 2\ell$ . Moreover, the same argument leading to (4.3) implies that

$$E_0 \ge \widetilde{E}_N(\boldsymbol{x}) \ge (\chi - 1)\log_{-}(2\ell) + C,$$

for all  $x \in \mathbb{R}_w^N$  such that  $E_0 \geq \widetilde{E}_N(x)$ , where  $C := C_1 + C_2$ . Therefore, by a direct method of calculus of variations, we get that  $\widetilde{E}_N$  has a global minimiser, which ends proof.

Remark 4.4. Similarly to Proposition 4.3 the proof above is given only for the discrete case, but is easily adaptable to the p-approximated one if  $p \ge 1$ , by Proposition 2.7(5) with s = 2.

Remark 4.5. The assumptions on V of Propositions 4.1 and 4.3 are in particular satisfied for V symmetric, coercive and convex. Propositions 4.1 and 4.3 are also true for any W satisfying Hypothesis 1 and bounded from below, with no required constraint on  $\chi$ , in particular for the linear Fokker-Planck equation, that is with  $V(x) = \frac{x^2}{2}$  and W = 0.

Let us now turn to the supercritical case. In the unconfined continuum modified Keller-Segel equation, it is known that solutions blow up in finite time if  $\chi > 1$ . The proof of non-existence of global-in-time solutions is obtained by computing the evolution of the second moment  $M_2(t)$  of solutions  $\rho(t)$  at any time t > 0. Then, a formal computation leads to

$$\frac{\mathrm{d}M_2}{\mathrm{d}t}(t) = \frac{\mathrm{d}}{\mathrm{d}t} \int_{\mathbb{R}} x^2 \,\mathrm{d}\rho(t,x) = 2(1-\chi).$$

Therefore the evolution of the second moment is linear in time with slope  $2(1 - \chi)$ . This slope is negative if  $\chi > 1$ , which implies that  $M_2(t)$  becomes zero in finite time leading to concentration of mass in finite time and contradiction with the assumption of global existence. We want here to show that our p-approximated discrete gradient flow (2.10) preserves this finite-time blow-up property for some numerical critical parameter  $\chi_2(N)$ , at least when all particles have same weight. Recall that, at the discrete level, we define blow-up as being the event of two particles colliding.

**Proposition 4.6.** Let p > 0 and consider the p-approximated discrete unconfined Keller-Segel gradient flow with  $w_i = 1/N$  for all  $i \in \{1, ..., N\}$ , on the whole time line  $[0, \infty)$ . All solutions blow up in finite time if  $\chi$  is greater than

$$\chi_2(N) := 1 + \frac{1}{N - 1}.$$

*Proof.* Let  $\boldsymbol{x}$  be a p-approximated discrete Keller-Segel gradient flow solution, emanating from an initial condition  $\boldsymbol{x^0} \in \mathbb{R}_w^N$ , defined on some maximal interval of existence  $[0, T^*)$ . Let us compute the evolution of the second moment of  $\mu_N$ , the empirical measure associated to  $\boldsymbol{x}$ , at any  $t \in [0, T^*)$ :

(4.4) 
$$\frac{dM_2}{dt}(t) = \frac{d}{dt} \int_{\mathbb{R}} x^2 d\mu_N(t, x) = \frac{d}{dt} \frac{1}{N} \sum_{i=1}^N x_i^2(t) = \frac{2}{N} \sum_{i=1}^N x_i(t) \frac{dx_i}{dt}(t).$$

In the following we drop the dependencies on time for the sake of simplicity.

Suppose  $T^* = \infty$ . We want to find a contradiction if  $\chi > \chi_2(N)$  by computing explicitly the evolution of the second moment in (4.4). Write  $\Delta_i^j := 1 + (\Delta x_j/\Delta x_i)^p$  for all  $i, j \in \{1, ..., N\}$  with  $i \neq j$ , recalling the convention  $\Delta x_1 = \Delta x_{N+1} = +\infty$  and setting  $\Delta_0^1 = \Delta_{N+2}^{N+1} = +\infty$ . By (2.10),

$$\sum_{i=1}^{N} x_i \frac{\mathrm{d}x_i}{\mathrm{d}t} = \underbrace{\sum_{i=1}^{N} \left( \left( \frac{x_i/\Delta x_i}{\Delta_{i-1}^i} + \frac{x_i/\Delta x_i}{\Delta_{i+1}^i} \right) - \left( \frac{x_i/\Delta x_{i+1}}{\Delta_{i}^{i+1}} + \frac{x_i/\Delta x_{i+1}}{\Delta_{i+2}^{i+1}} \right) \right)}_{i}$$

$$-\frac{2\chi}{N} \sum_{i=1}^{N} x_i \sum_{\substack{j=1\\j\neq i}}^{N} W'(x_i - x_j).$$

First, compute  $A_1$  by appropriately rearranging the sum-terms,

$$A_{1} = \sum_{i=3}^{N-2} \left( \left( \frac{x_{i}/\Delta x_{i}}{\Delta_{i-1}^{i}} + \frac{x_{i}/\Delta x_{i}}{\Delta_{i+1}^{i}} \right) - \left( \frac{x_{i}/\Delta x_{i+1}}{\Delta_{i}^{i+1}} + \frac{x_{i}/\Delta x_{i+1}}{\Delta_{i+2}^{i+1}} \right) \right) + \left( 1 + \frac{1}{\Delta_{3}^{2}} \right) + \left( 1 + \frac{1}{\Delta_{N-1}^{N}} \right) - \left( \frac{x_{2}/\Delta x_{3}}{\Delta_{2}^{3}} + \frac{x_{2}/\Delta x_{3}}{\Delta_{4}^{3}} \right) + \left( \frac{x_{N-1}/\Delta x_{N-1}}{\Delta_{N}^{N-1}} + \frac{x_{N-1}/\Delta x_{N-1}}{\Delta_{N-2}^{N-1}} \right).$$

Also,

$$\begin{split} \sum_{i=3}^{N-2} \left( \left( \frac{x_i/\Delta x_i}{\Delta_{i-1}^i} + \frac{x_i/\Delta x_i}{\Delta_{i+1}^i} \right) - \left( \frac{x_i/\Delta x_{i+1}}{\Delta_{i}^{i+1}} + \frac{x_i/\Delta x_{i+1}}{\Delta_{i+2}^{i+1}} \right) \right) \\ &= \left( \frac{x_3/\Delta x_3}{\Delta_{4}^3} + \frac{x_3/\Delta x_3}{\Delta_{2}^3} \right) - \left( \frac{x_{N-2}/\Delta x_{N-1}}{\Delta_{N-2}^{N-1}} + \frac{x_{N-2}/\Delta x_{N-1}}{\Delta_{N}^{N-1}} \right) + \sum_{i=3}^{N-2} \left( \frac{1}{\Delta_{i}^{i+1}} + \frac{1}{\Delta_{i+2}^{i+1}} \right) \\ &+ \left( 1 + \frac{1}{\Delta_{3}^2} \right) + \left( \frac{1}{\Delta_{4}^3} + \frac{1}{\Delta_{2}^3} \right) + \left( \frac{1}{\Delta_{N-1}^{N-1}} + \frac{1}{\Delta_{N-2}^{N-1}} \right) + \left( 1 + \frac{1}{\Delta_{N-1}^{N}} \right). \end{split}$$

Hence, by combining the last two computations, we get

$$\begin{split} A_1 &= \sum_{i=3}^{N-2} \left( \frac{1}{\Delta_i^{i+1}} + \frac{1}{\Delta_{i+2}^{i+1}} \right) + \left( \frac{1}{\Delta_4^3} + \frac{1}{\Delta_2^3} \right) + \left( \frac{1}{\Delta_{N-2}^{N-1}} + \frac{1}{\Delta_N^{N-1}} \right) \\ &+ \left( 1 + \frac{1}{\Delta_2^3} \right) + \left( 1 + \frac{1}{\Delta_{N-1}^N} \right) = \sum_{i=2}^{N-1} \left( \frac{1}{\Delta_i^{i+1}} + \frac{1}{\Delta_{i+1}^i} \right) + 2 = \sum_{i=2}^{N-1} 1 + 2 = N. \end{split}$$

Then, compute  $A_2$  by using the anti-symmetry of  $W'(x) = \frac{1}{x}$ ,

$$A_{2} = (x_{1} - x_{2})W'(x_{1} - x_{2}) + (x_{1} - x_{3})W'(x_{1} - x_{3}) + \dots + (x_{1} - x_{N})W'(x_{1} - x_{N}) + (x_{2} - x_{3})W'(x_{2} - x_{3}) + (x_{2} - x_{4})W'(x_{2} - x_{4}) + \dots + (x_{2} - x_{N})W'(x_{2} - x_{N}) + \dots + (x_{N-1} - x_{N})W'(x_{N-1} - x_{N}) = (N-1) + (N-2) + \dots + 1 = \sum_{i=1}^{N-1} (N-i) = \frac{(N-1)N}{2}$$

Therefore, for all  $t \in [0, T^*) = [0, \infty)$ ,

$$\frac{\mathrm{d}M_2}{\mathrm{d}t}(t) = \frac{2}{N} \left( N - \frac{2\chi}{N} \frac{(N-1)N}{2} \right) = 2\left( 1 - \chi \left( 1 - \frac{1}{N} \right) \right) = 2\left( 1 - \frac{1}{N} \right) (\chi_2(N) - \chi).$$

Hence the evolution of the second moment is linear with a negative slope, since by assumption  $\chi > \chi_2(N)$ , which clearly contradicts the fact that the maximal time of existence  $T^* = \infty$ , and therefore the solution  $\boldsymbol{x}$  exists only up to a finite time:  $T^* < \infty$ . At exactly that time, only two things may happen: either the norm of the solution equals  $+\infty$ , i.e.,  $|\boldsymbol{x}|_w = +\infty$ , or two or more particles collide. The first option is not plausible since trivially the second moment of an empirical measure is finite at all times. We are thus only left with the collision of particles, that is  $\boldsymbol{x}$  has to blow up in finite time.

4.1.2. Simulations. We give here a few simulations for the modified Keller-Segel equation showing various blow-up characteristics when V=0. As we want to capture the blow-up we used an adaptive time-step size as follows. For every time step  $n \in \{0, ..., M-1\}$ , suppose we have a time-step size  $\Delta_n t$ , and compute the velocity  $v_i^n$  of each particle  $x_i^n$ ,  $1 \le i \le N$ . Then fix a tolerance  $\delta = 0.25$ , and define

$$\delta_i = \begin{cases} \Delta_n t & \text{if } \Delta_n t \leq \frac{\delta \min_p(\Delta x_i^n, \Delta x_{i+1}^n)}{|v_i^n|}, \\ \frac{\delta \min_p(\Delta x_i^n, \Delta x_{i+1}^n)}{|v_i^n|} & \text{otherwise.} \end{cases}$$

Finally, renew  $\Delta_n t := \min_{i \in \{1,\dots,N\}} \delta_i$ , compute the positions  $x_i^{n+1}$  with the new  $\Delta_n t$ , and start over until  $\Delta_n t > 10^{-7}$ , in which case stop the simulation. Note that we took  $\Delta_0 t = 10^{-5}$  as the very initial time-step size. In Figures 7, 8 and 9 the simulations shown stopped due to this adaptive time-step procedure.

Figures 7 and 8 show the results of simulations with initial continuum profile  $\rho_0^{\text{heat}}$ , with  $I_{\text{init}} = [-2.5, 2.5]$ . From Figure 7 one can see that the scheme we used captures nicely the formation of the blow-up for a supercritical parameter  $\chi$ .

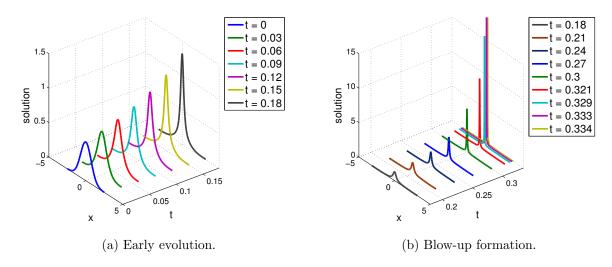
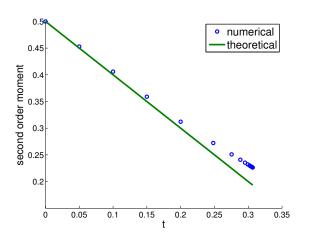
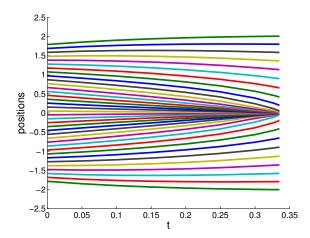


FIGURE 7. The Keller-Segel equation with  $\chi = 1.5$  for N = 50.

Figure 8a shows that the evolution of the second moment is linear for some range of time, with a slope that deviates slightly from the theoretical one, as expected from Proposition 4.6 and its proof. For larger times, as the blow-up is approached and the time step refined, the numerical slope deviates even more from the theoretical one.

Figure 8b shows the trajectories of the particles up to the first blow-up time, defined numerically as the first time when the distance between two particles gets smaller than some chosen number  $d_{\min}$ , or equivalently when the adaptive time-step size  $\Delta_n t$  gets larger than  $10^{-7}$ , see above. A possible procedure to continue the evolution after the first blow-up time is by merging particles into a new heavier one whenever the distance between these particles gets less than a certain threshold, say proportional to  $d_{\min}$ ; the weight of the new particle is then chosen to be the sum of the merged particles, and the position the barycentre of the merged particles. This procedure might give us an idea of how the particles behave after the first blow-up, but we found that it is not very accurate since the post-collision trajectories strongly depended on the choice of the threshold, which is arbitrary. We therefore found that the analysis of the post-collision behaviour is very delicate without having clear criteria for deciding when to merge particles and how many simultaneously. We thus leave this issue for further analysis and future work.



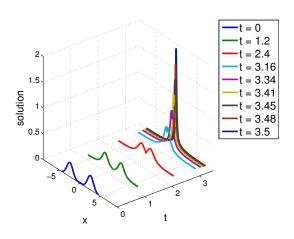


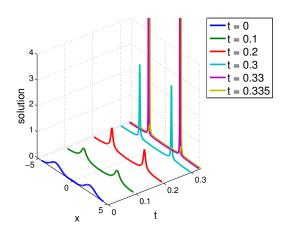
- (a) Evolution of the second moment for N = 100.
- (b) Particle trajectories up to first numerical blow-up for N=50 (not all particles are represented).

FIGURE 8. The blow-up of the Keller-Segel equation with  $\chi = 1.5$ .

Figure 9 shows the result with a continuum two-bump initial profile, with  $I_{\text{init}} = [-4.5, 4.5]$ :

$$\rho_0(x) = \frac{1}{2\sqrt{4\pi t_0}} e^{-\frac{(x+2)^2}{4t_0}} + \frac{1}{2\sqrt{4\pi t_0}} e^{-\frac{(x-2)^2}{4t_0}} \quad \text{with } t_0 = 0.25.$$





- (a) Evolution with  $\chi = 1.8$  for N = 50.
- (b) Evolution with  $\chi = 3$  for N = 100.

FIGURE 9. Blow-up formation for the Keller-Segel equation with two initial Gaussian bumps.

Figure 9 shows the possible formation of several Dirac masses, according to how much attraction is involved in the system and to how many bumps are present at the beginning. It seems like the more attraction, the more Dirac masses can form. Note that the two peaks in Figure 9b are actually of same height, despite a displaying artifact; also, even if not clear from Figure 9b, the two peaks get slightly closer to each other with time, but then blow up before merging.

4.2. The Keller-Segel equation with nonlinear diffusion. Let us now consider the onedimensional Keller-Segel equation with nonlinear diffusion, that is the continuum gradient flow (1.2) with  $H(\rho) = \frac{\rho^m}{m-1}$ , m > 1, V = 0 and  $W(x) = 2\chi \log |x|$  (and W(0) := 0). The initial continuum profile we used here is  $\rho_0^{\text{heat}}$ , with  $I_{\text{init}} = [-2.5, 2.5]$ .

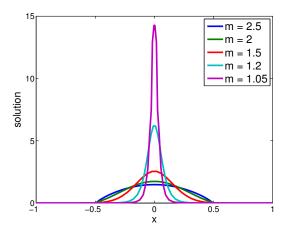


FIGURE 10. The nonlinear Keller-Segel equation with  $\chi = 1.4$  for different choices of m and N = 50 at T = 4, with  $\Delta t = 10^{-5}$ .

Each curve in Figure 10 is a good approximation of a steady state for the Keller-Segel equation with nonlinear diffusion. For each m, the steady state is different; and as m tends to 1 the steady state "squeezes" and looks as if it is approaching a Dirac mass, which is the "steady state" of the Keller-Segel equation with linear diffusion studied in Section 4.1.

4.3. A compactly supported potential with nonlinear diffusion. We consider here the continuum gradient flow when  $H(\rho) = \frac{\rho^m}{m-1}$ , m > 1, V = 0 and  $W(x) = -c \max(1-|x|,0)+c$ , c > 0, is a compactly supported interaction potential. In Figure 11 the considered continuum initial profile is a uniform distribution on the interval [-2,2]. Here  $I_{\text{init}} = [-2,2]$  and the extreme particles were set to have weights equal to 0.001.

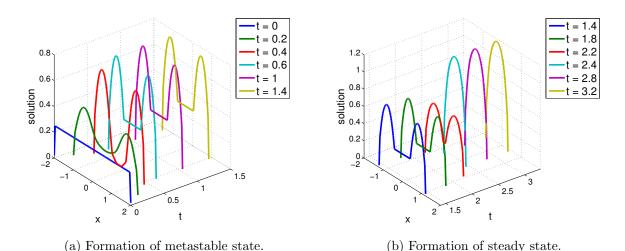


FIGURE 11. Compactly supported potential  $W(x) = -c \max(1 - |x|, 0) + c$  with nonlinear diffusion with c = 8 and m = 3, for N = 80, with  $\Delta t = 10^{-5}$ .

Figure 11a shows the formation of a metastable state made of two bumps, while Figure 11b shows how this metastable state breaks into a single-bumped steady state. This behaviour, for exactly

this interaction potential (up to a multiplicative constant, c), was already noted in [12, Example 3] when using a finite-volume scheme on gradient flows. The bumps are actually supposed to be disconnected since no particles were found numerically in between these bumps, although this does not seem to be the case on the plots. This is because the particle of each bump which is closest to the origin is not located at the actual boundary of the exact solution's corresponding bump, and so the line connecting the two bumps is not at 0 but higher.

Appendix A. Convergence of the p-approximated gradient flow in one dimension

In this appendix we give the proof of the convergence of the p-approximated gradient flow (2.10) to the discrete one (2.7) in one dimension, see Section 2.3. Before doing so we need to recall some notions and results from monotone operator theory.

If  $A: \{y \in X \mid Ay \neq \emptyset\} \subset X \to X^*$ , where X is a Banach space and  $X^*$  its dual, is a maximal monotone operator, then the graph of A is given by graph $(A) = \{(x, Ax) \mid x \in \{y \in X \mid Ay \neq \emptyset\}\}$ .

**Definition A.1** (Graph convergence). Let X be a reflexive Banach space,  $(A_k)_{k\in\mathbb{N}}$  a sequence of maximal monotone operators from X to  $X^*$ , and A a maximal monotone operator from X to  $X^*$ . We say that  $(A_k)_{k\in\mathbb{N}}$  graph-converges to A if, for every  $(x,y)\in \operatorname{graph}(A)$ , there exists a sequence  $((x_k,y_k))_{k\in\mathbb{N}}$  with  $(x_k,y_k)\in\operatorname{graph}(A_k)$  for all  $k\in\mathbb{N}$  such that  $x_k\to x$  strongly in X and  $y_k\to y$  strongly in  $X^*$  as  $x_k\to\infty$ .

Remark A.2. One can check that if  $\phi$  is a lower semi-continuous, convex function from a reflexive Banach space to  $\mathbb{R}$ , then  $\partial \phi$  is a maximal monotone operator, see [3, Section 3.7.1].

For the sake of completeness we give the defintion of  $\Gamma$ -convergence.

**Definition A.3** ( $\Gamma$ -convergence). Let  $(X_k)_{k\in\mathbb{N}}$  be a sequence of metric spaces endowed with a distance d and  $(\phi_k)_{k\in\mathbb{N}}$  be a sequence of functionals  $\phi_k \colon X_k \to \mathbb{R}$  for all  $k \in \mathbb{N}$ . We say that  $(\phi_k)_{k\in\mathbb{N}}$   $\Gamma$ -converges to  $\phi$  if the following two conditions are met for all  $u \in X$ :

- (i) ("liminf" condition)  $\phi(u) \leq \liminf_{k \to \infty} \phi_k(u_k)$  for all sequences  $(u_k)_{k \in \mathbb{N}}$  with  $u_k \in X_k$  for all  $k \in \mathbb{N}$  and  $d(u_k, u) \to 0$  as  $k \to \infty$ ,
- (ii) ("limsup" condition)  $\limsup_{k\to\infty} \phi_k(u_k) \leq \phi(u)$  for some sequence  $(u_k)_{k\in\mathbb{N}}$  with  $u_k\in X_k$  for all  $k\in\mathbb{N}$  and  $d(u_k,u)\to 0$  as  $k\to\infty$ .

Theorem A.4 connects the notion of  $\Gamma$ -convergence to that of graph-convergence in finite dimension. It is a consequence of [3, Theorem 3.66] and the fact that  $\Gamma$ -convergence is equivalent to Mosco convergence in finite dimension. Indeed, in a general dimensional setting, Mosco convergence means that the "liminf" condition of  $\Gamma$ -convergence holds for the weak topology and the "limsup" condition holds for the strong topology, see [26, Definition 2.2] and [3, Definition 3.17].

**Theorem A.4.** Let X be a finite-dimensional Banach space. Let  $(\phi_k)_{k\in\mathbb{N}}$  be a sequence of lower semi-continuous, convex functions with  $\phi_k \colon X \to \mathbb{R}$  for all  $k \in \mathbb{N}$ , and  $\phi \colon X \to \mathbb{R}$  a lower semi-continuous, convex function. If  $(\phi_k)_{k\in\mathbb{N}}$   $\Gamma$ -converges to  $\phi$ , then  $(\partial \phi_k)_{k\in\mathbb{N}}$  graph-converges to  $\partial \phi$ .

We can now give the general convergence and regularity result, whose proof can be deduced by [3, Theorem 3.74], [9, Theorem 3.1] and [4, Theorem 1, Section 3.2].

**Theorem A.5.** Let X be a Hilbert space. Let  $(\phi_k)_{k\in\mathbb{N}}$  be a sequence of lower semi-continuous, convex functions with  $\phi_k \colon X \to \mathbb{R}$  for all  $k \in \mathbb{N}$  and  $\phi \colon X \to \mathbb{R}$  be a lower semi-continuous, convex function. Suppose that  $(\partial \phi_k)_{k\in\mathbb{N}}$  graph-converges to  $\partial \phi$ . Consider the following differential inclusions, for all  $k \in \mathbb{N}$ .

$$u'_k(t) \in -\partial \phi_k(u_k(t)), \quad u_k(0) = u_k^0 \quad \text{for almost every } t \in (0, T],$$

and

$$u'(t) \in -\partial \phi(u(t)), \quad u(0) = u_0 \quad \text{for almost every } t \in (0, T],$$

where  $u, u_k : [0, T] \to X$  are the unknown curves. Unique solutions u and  $u_k$  exist and

- (1)  $u_k$  and u are continuous on [0,T],
- (2)  $u'_k$  and u' are right-continuous on [0,T].

Moreover, assume that  $u_k^0 \to u_0$  strongly and that in this case,  $\phi_k(u_k^0) \to \phi(u_0)$  as  $k \to \infty$ . Then

- (3)  $u_k \to u$  uniformly on [0,T] as  $k \to \infty$ ,
- (4)  $\int_0^T t |u_k'(t) u'(t)|^2 dt \to 0 \text{ as } k \to \infty,$
- $(5) \ u_k' \rightarrow u' \ strongly \ in \ L^2([0,T],X) \ as \ k \rightarrow \infty \ (so \ u_k'(t) \rightarrow u'(t) \ for \ almost \ every \ t \in [0,T]),$
- (6)  $\phi_k(u_k) \to \phi(u)$  uniformly on [0,T] as  $k \to \infty$ ,

It is not hard to see that  $\widetilde{E}_N^p$ , see (2.11),  $\Gamma$ -converges to  $\widetilde{E}_N$ , see (2.4), as  $p \to \infty$ . Furthermore, it is easily verified that  $\widetilde{E}_N^p(\boldsymbol{x^0}) \to \widetilde{E}_N(\boldsymbol{x^0})$  as  $p \to \infty$ , where  $\boldsymbol{x^0} \in \mathbb{R}_w^N$  is taken here to be the initial condition for both the discrete and p-approximated discrete gradient flows (2.7) and (2.10). Therefore, in order to use Theorems A.4 and A.5 combined, we are only left with checking that  $\widetilde{E}_N^p$  and  $\widetilde{E}_N$  are lower semi-continuous and convex. The first condition is trivial to verify based on the assumptions on H, V and W, whereas the convexity condition is shown below whenever V and W are assumed to be convex. Actually the following proposition shows the convexity of  $\widetilde{E}_N$  only, but also holds for  $\widetilde{E}_N^p$  since the proof is easily adapted from the classical minimum function to the p-approximated one.

**Proposition A.6.** Let d = 1 and the confinement and interaction potentials V and W be convex. Then the discrete energy  $\widetilde{E}_N$  defined in (2.4) is convex.

*Proof.* Let  $\lambda \in [0,1]$  and  $\boldsymbol{x}, \boldsymbol{y} \in \mathbb{R}_w^N$ . Then, by the facts that min is concave on  $\mathbb{R}^2$  and h is non-increasing and convex on  $(0,\infty)$ , we know that  $h \circ \min$  is convex on  $(0,\infty)^2$ , where  $\circ$  is the composition operator. Define, for all  $a, b \in \mathbb{R}$ ,  $[a,b]_{\lambda} = \lambda a + (1-\lambda)b$ , and  $r_i(\boldsymbol{x}) = \min(\Delta x_i, \Delta x_{i+1})$  and  $r_i(\boldsymbol{y}) = \min(\Delta y_i, \Delta y_{i+1})$ . Therefore, since V and W are convex,

$$\begin{split} \widetilde{E}_{N}([\boldsymbol{x},\boldsymbol{y}]_{\lambda}) &= \sum_{i=1}^{N} w_{i} h\left(\frac{1}{w_{i}} \min([\Delta x_{i}, \Delta y_{i}]_{\lambda}, [\Delta x_{i+1}, \Delta y_{i+1}]_{\lambda})\right) \\ &+ \sum_{i=1}^{N} w_{i} V([x_{i}, y_{i}]_{\lambda}) + \frac{1}{2} \sum_{i=1}^{N} \sum_{\substack{j=1 \ j \neq i}}^{N} w_{i} w_{j} W([x_{i} - x_{j}, y_{i} - y_{j}]_{\lambda}) \\ &\leq \lambda \sum_{i=1}^{N} w_{i} h\left(\frac{r_{i}(\boldsymbol{x})}{w_{i}}\right) + (1 - \lambda) \sum_{i=1}^{N} w_{i} h\left(\frac{r_{i}(\boldsymbol{y})}{w_{i}}\right) + \lambda \sum_{i=1}^{N} w_{i} V(x_{i}) + (1 - \lambda) \sum_{i=1}^{N} w_{i} V(y_{i}) \\ &+ \frac{\lambda}{2} \sum_{i=1}^{N} \sum_{\substack{j=1 \ j \neq i}}^{N} w_{i} w_{j} W(x_{i} - x_{j}) + \frac{1 - \lambda}{2} \sum_{i=1}^{N} \sum_{\substack{j=1 \ j \neq i}}^{N} w_{i} w_{j} W(y_{i} - y_{j}) \\ &= \lambda \widetilde{E}_{N}(\boldsymbol{x}) + (1 - \lambda) \widetilde{E}_{N}(\boldsymbol{y}). \end{split}$$

Hence convexity of  $\widetilde{E}_N$ .

Theorem A.4 now tells us that  $\partial_w \widetilde{E}_N^p$  graph-converges to  $\partial_w \widetilde{E}_N$  as  $p \to \infty$ , and Theorem A.5 tells us in which sense the *p*-approximated discrete gradient flow converges to the discrete one and also gives us some regularity on the discrete and *p*-approximated discrete gradient flow solutions. The use of the *p*-approximated gradient flow (2.10) is therefore justified to approximate (2.7) (at least in the case when *V* and *W* are convex and d = 1).

**Acknowledgements.** JAC, YH and FSP are supported by Engineering and Physical Sciences Research Council grant EP/K008404/1. JAC is also supported by the project MTM2011-27739-C04-02 DGI (Spain) and by the Royal Society through a Wolfson Research Merit Award (JAC). GW is supported by ISF grant 998/5.

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